


Astrophysical Imprints of Wakefields: NS-NS Collision, γ -emissions from Blazars, Pinpointed High Energy Cosmic Rays




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Nature's wakefield accelerators in cosmos

1. Collision of neutron star - neutron star
 2. Episodic eruption of γ -emissions from Blazars
 3. Pinpointed high energy cosmic rays (and neutrinos)
- 

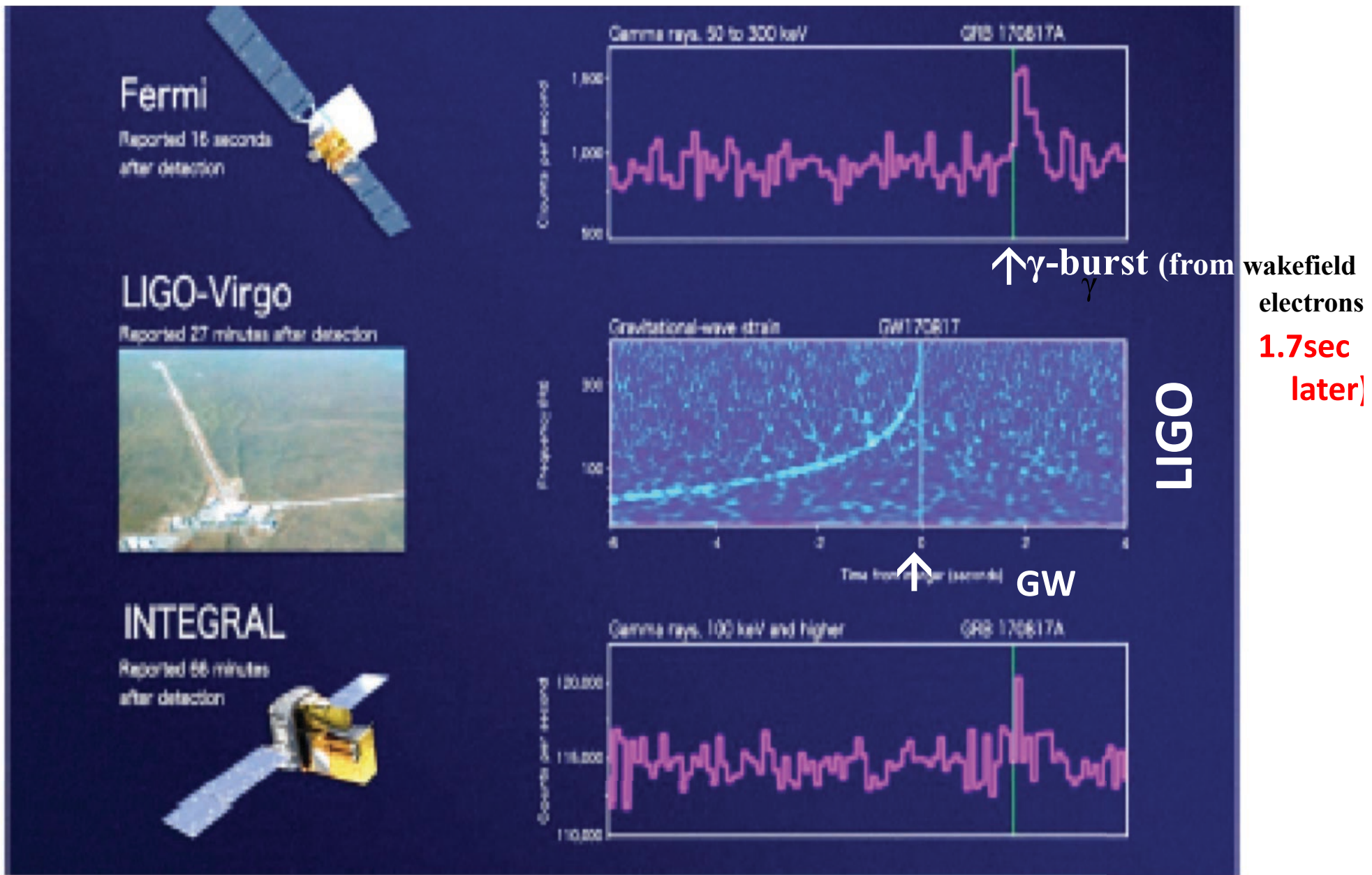


Fig. 5. Gamma-ray emission detected by Fermi and Integral satellites from the neutron star merging event (GW178017) delayed by 1.7 seconds compared with gravitational wave burst [79]. This time difference may be explained by the time to build-up the system for the acceleration of charged particles, described in the present

Barry Barish: 2017 Nobel Observation of Gravitational Waves



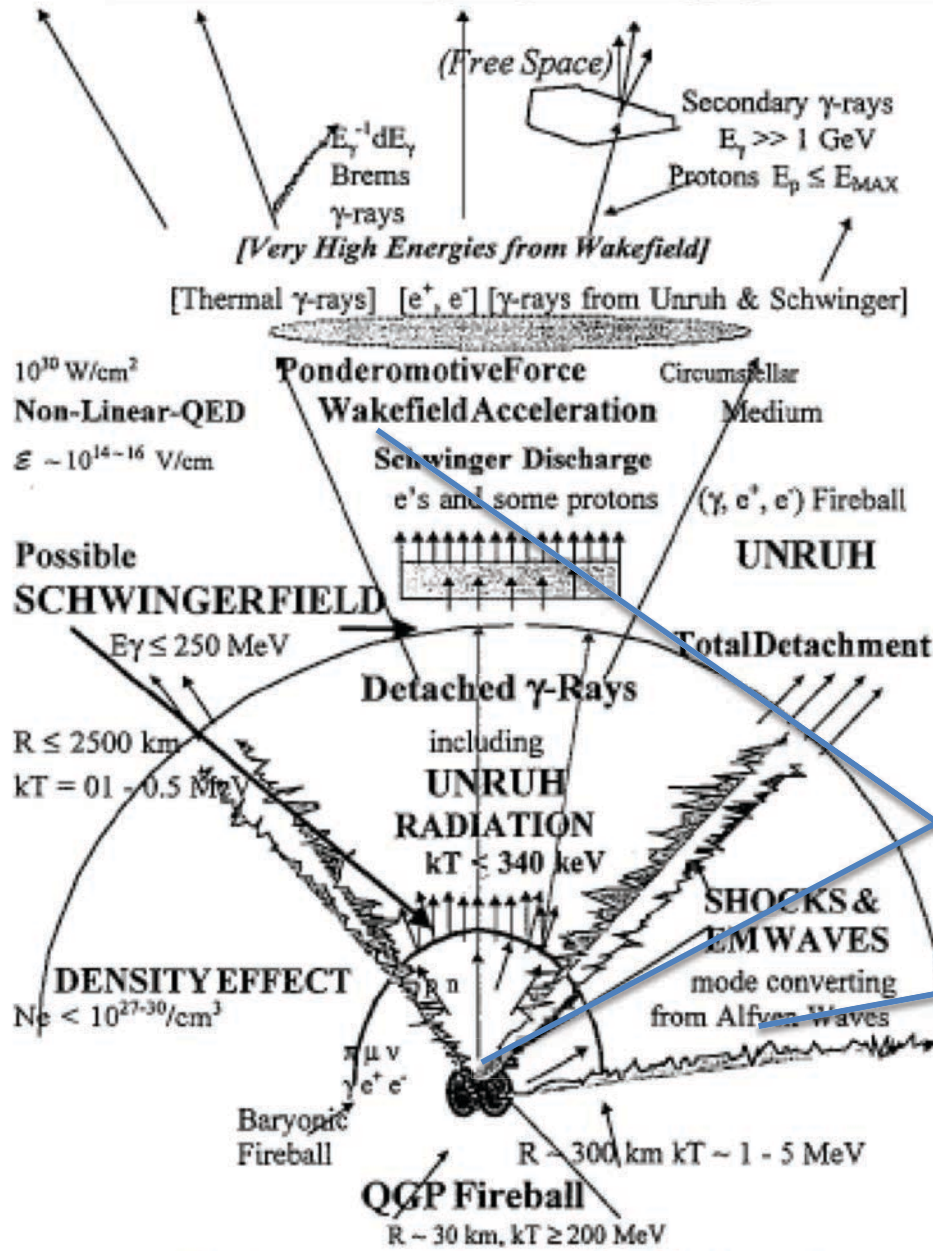
T. Ebisuzaki

With Professor B. Barish at **LIGO**, Caltech

GRB including high energy particles

**Prophetic picture
(2000)**

**NS-NS collision
triggers →**



- QGP (Quark-Gluon plasma)
- Shocks / **gravitational waves**
- Accretion disk
- Jets
- Alfvén waves and EM waves
- Wakefield acceleration
- GRB (**gamma bursts**)

.....

Figure 8. A schematic illustration of the proposed concept.

Takahashi, Y., Hillman, L.W., Tajima, T., in High Field Science, (Kluwer, NY, 2000) p.171.

Spacetime scales of NS-NS collision

Accretion disk

Jets/

Alfven waves and EM waves/

Wakefield acceleration / $3 \times 10^5 \text{ km}$

GRB (gamma ray bursts) $t = 1 \text{ s}$

Unruh radiation 3000 km $t = 10 \text{ ms}$

Baryon fireball
 300 km $t = 1 \text{ ms}$

Shocks / gravitational waves
QGP (Quark-Gluon plasma)
 30 km $t = 0$

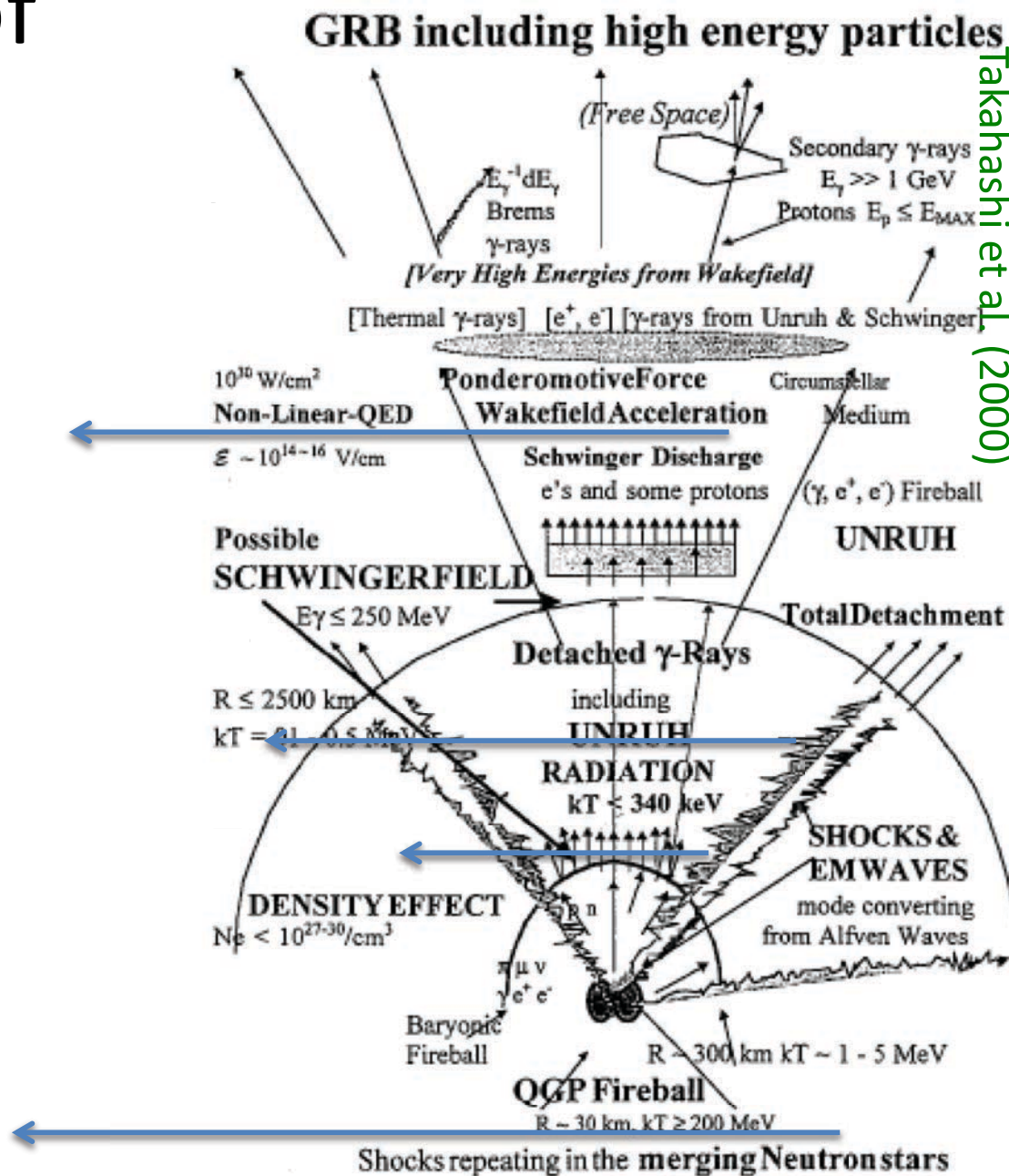
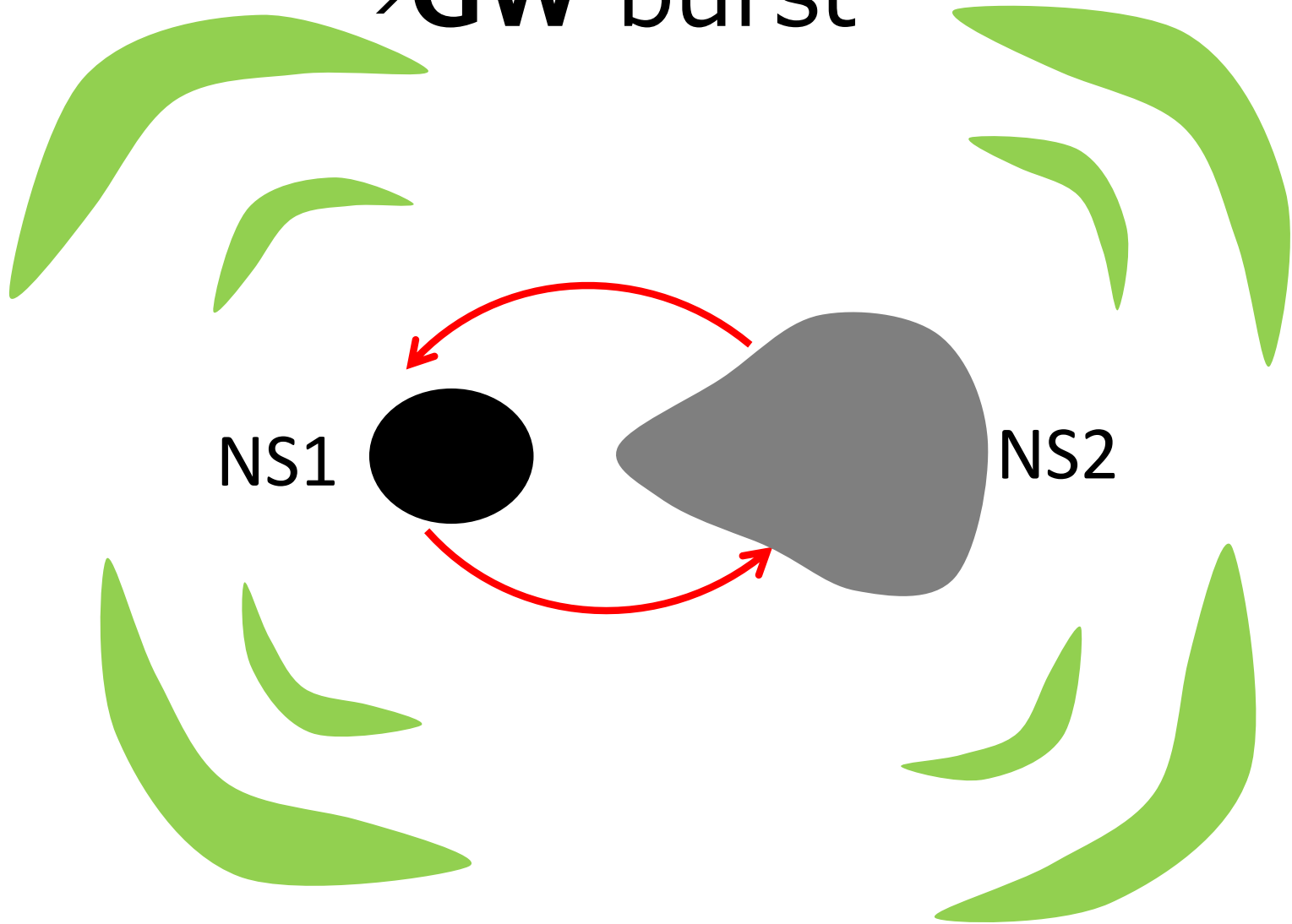


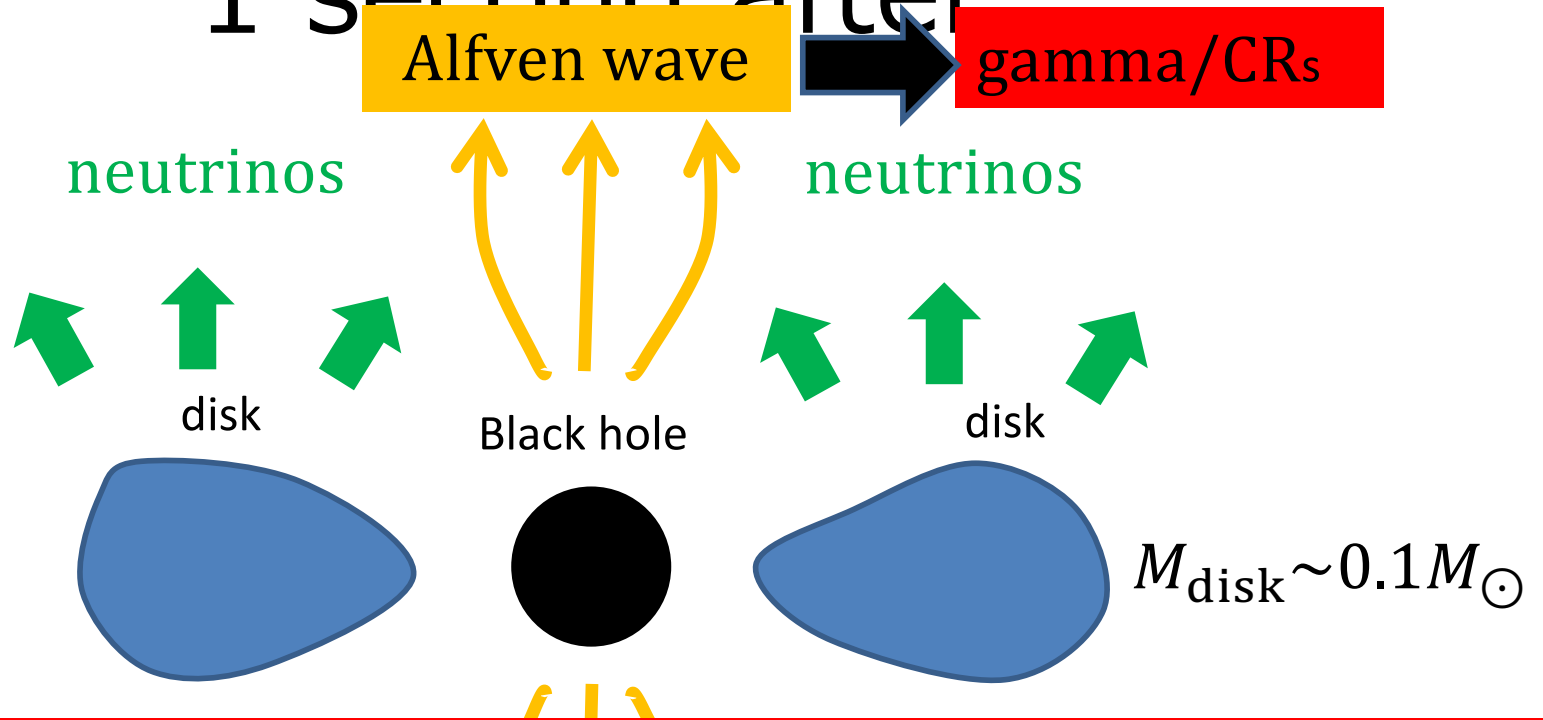
Figure 8. A schematic illustration of the proposed concept.

NS-NS merger → **GW** burst



NS-NS merger \rightarrow BH + Disk

1 second after

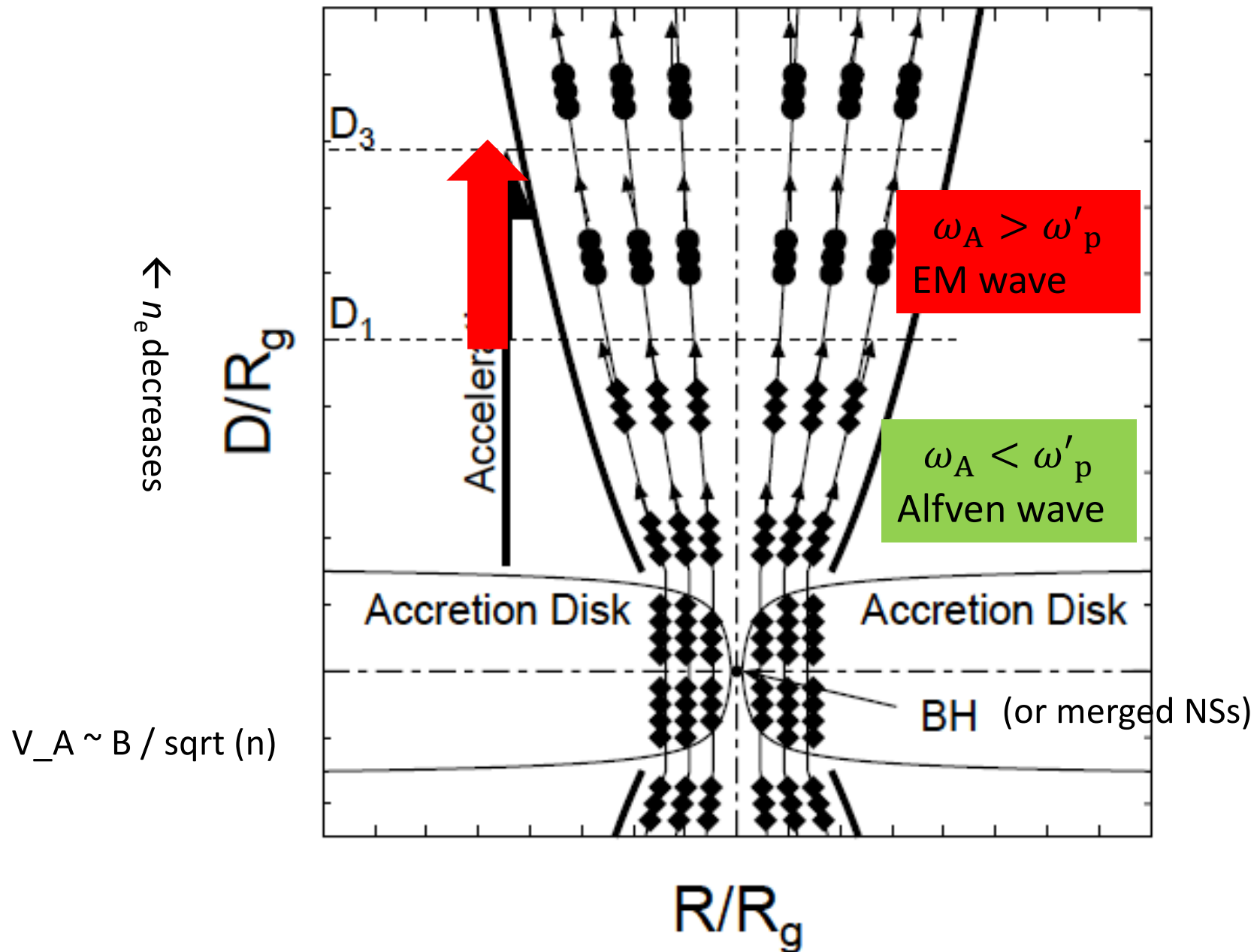


$$L_{\nu} \sim 10^{52} \text{ erg/s} \sim L_A$$

Central Engine of GRB/Hypernova

Alfven wave

Wakefield generation in Jet



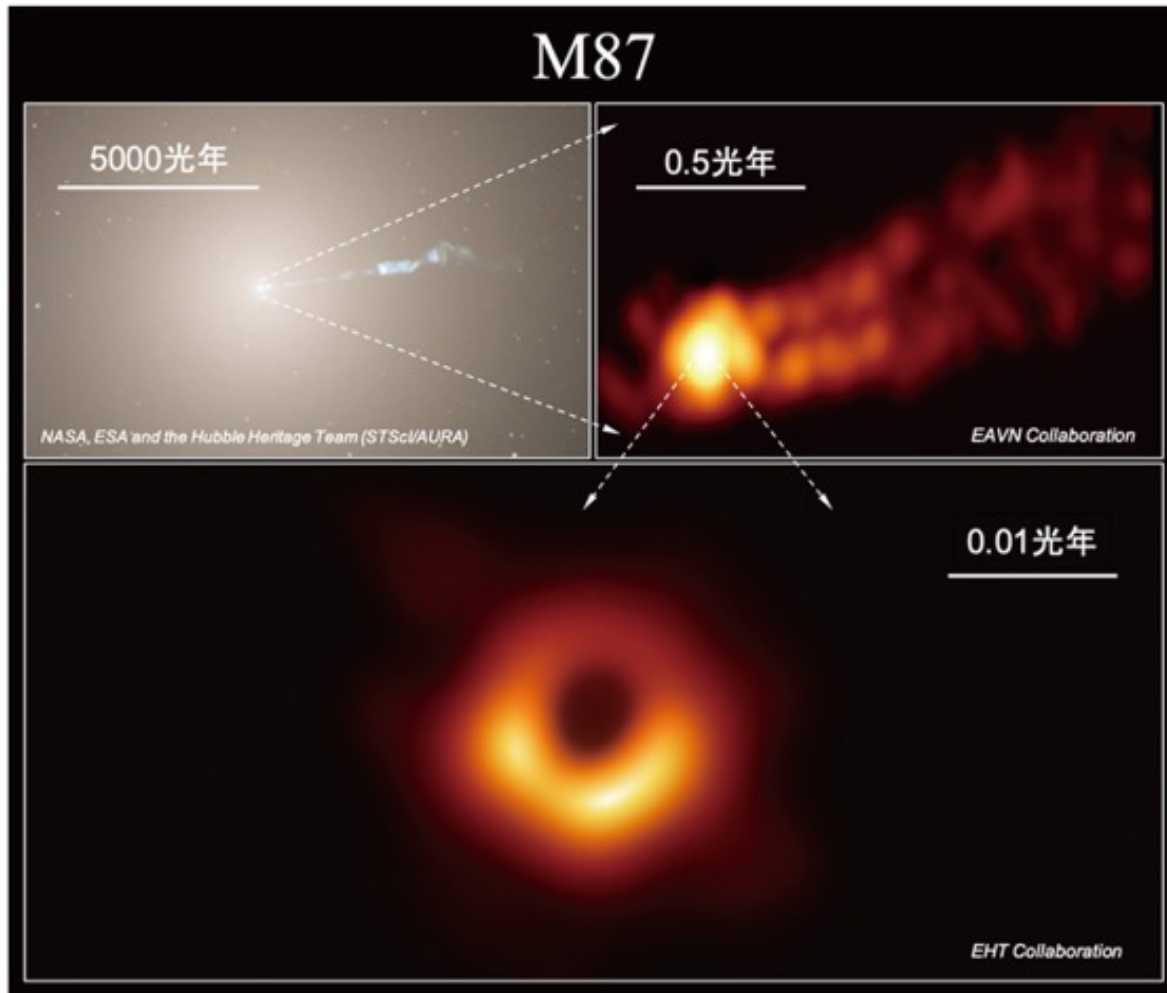
Nature's wakefield accelerators in cosmos



2. Episodic eruption of γ -emissions from Blazars

Discovery of **Blackhole** and **Prediction**

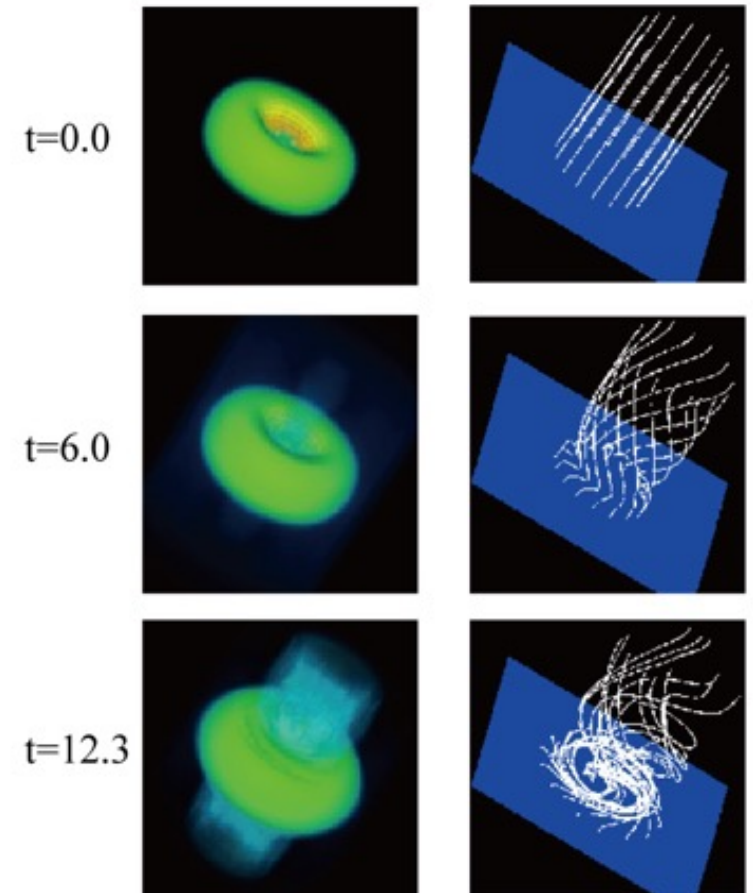
M87 blackhole: by Event Horizon Telescope (2019)



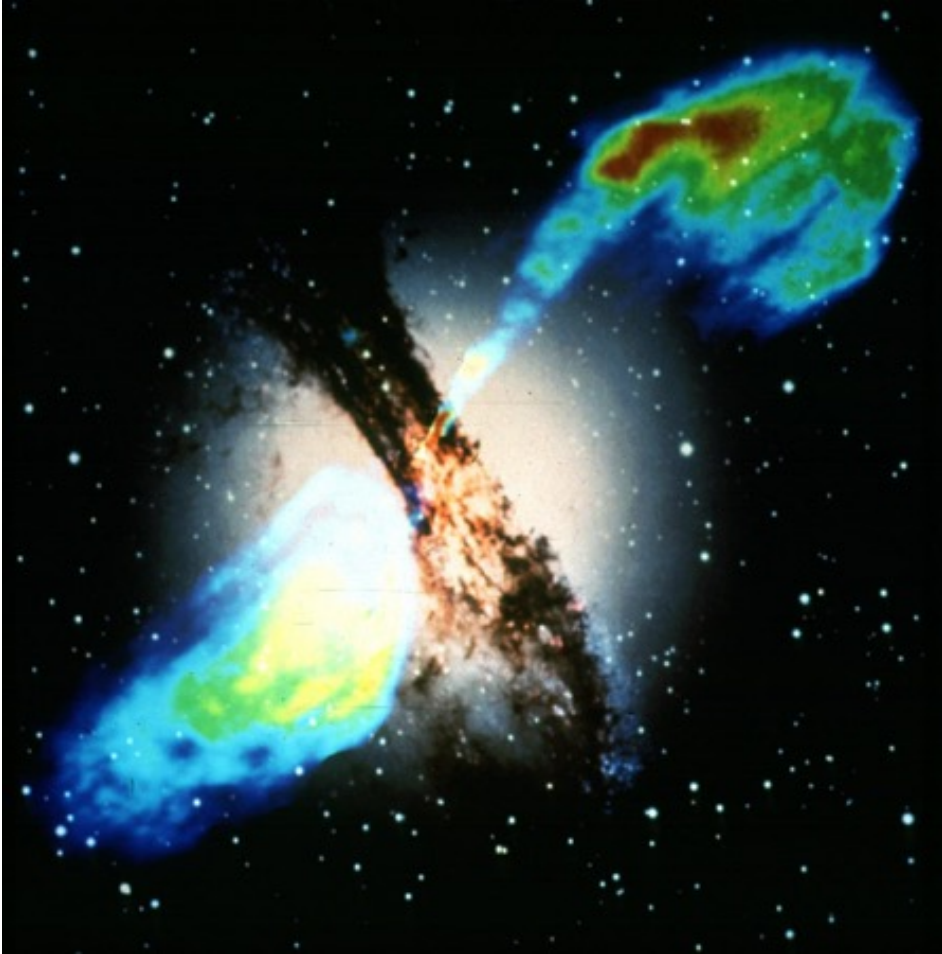
Suggestion: Tajima and Shibata
“**Plasma Astrophysics**”

(textbook, 1997)

3D Structure of Disk and Jet



“Physiology” of various AGNs



Cen A

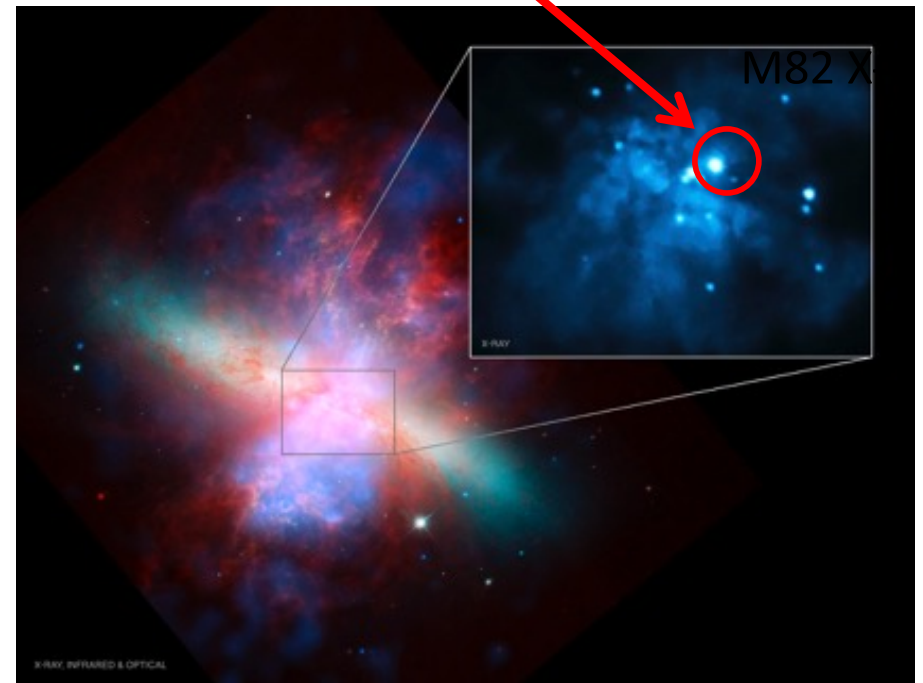
- Distance : 3.4Mpc
- **Radio Galaxy**
 - Nearest
 - **Brightest radio source**
- Elliptical Galaxy
- Black hole at the center w/
relativistic jets

M82: Nearest Starburst Galaxy

M82 X-1: 1000-10000 M_{\odot} BH



Just after the collision with M81

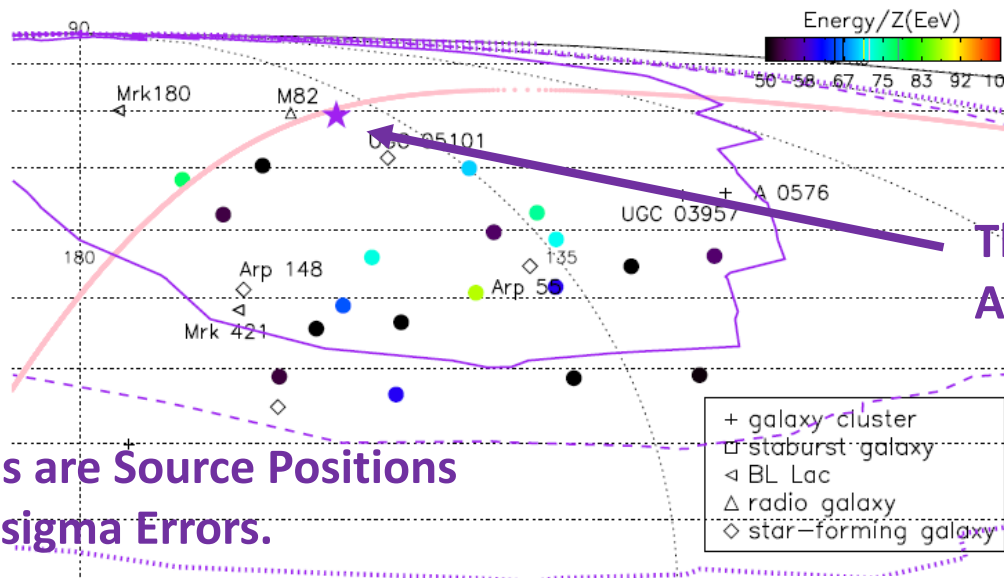


Composite of X-ray, IR, and optical emissions

NASA / CXC / JHU / D. Strickland; optical: NASA / ESA / STScI / AURA / Hubble Heritage Team; IR: NASA / JPL-Caltech / Univ. of AZ / C. Engelbracht; inset – NASA / CXC / Tsinghua University / H. Feng et al.

TA Hot Spot: UHECRs from M82?

He, Kusenko, Nagataki + PRD 2016.



The most likely Source Position
As a Result of Our Analysis.

Purple Lines are Source Positions
With 1,2,3-sigma Errors.

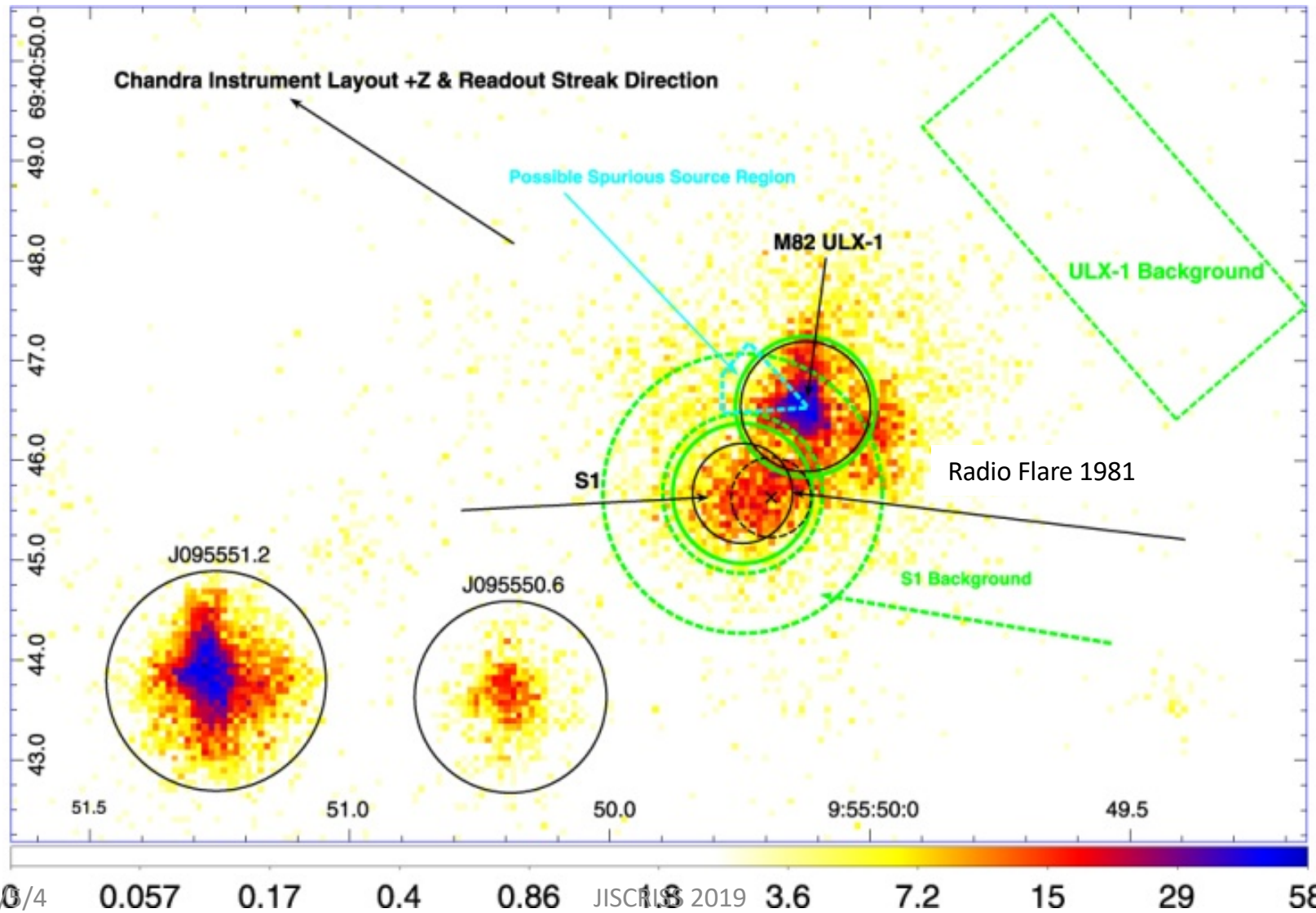
M82 is very Close
from the most likely
Source Position!

Source Name	Source Type	Distance (Mpc)	A_1 ($^\circ$)	A_2 ($^\circ$)	$P/P_{\text{bes-fit}}$ (%)
best-fit	-	-	$17.4^{+17.0}_{-11.0}$	$9.4^{+3.7}_{-0.3}$	100
M82	starburst galaxy	3.4	17.6	9.6	99.8
UGC 05101	star-forming galaxy	160.2	11.6	9.2	96.9
Mrk 180	blazar	185	19.9	9.3	91.3
UGC 03957	galaxy cluster	150.3	14.9	9.5	67.4
A 0576	galaxy cluster	169.0	17.0	9.4	63.4
Arp 55	star-forming Galaxy	162.7	1.9	9.7	55.3
Arp 148	star-forming Galaxy	143.3	10.5	10.0	41.8
Mrk 421	blazar	134	11.2	9.9	35.6

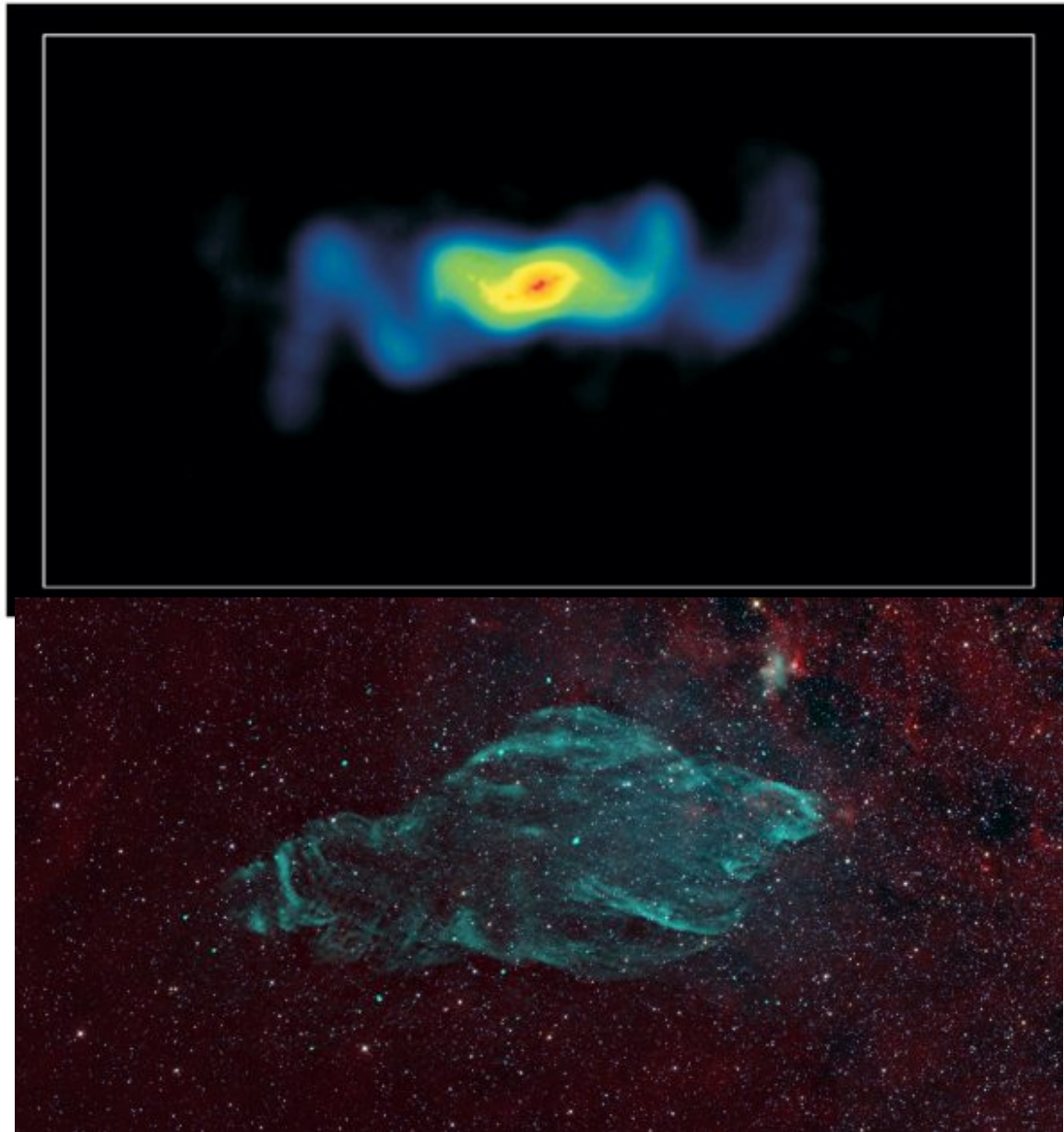
An AGN-like Jet in M82?

X-ray/Radio (flare in 1981)

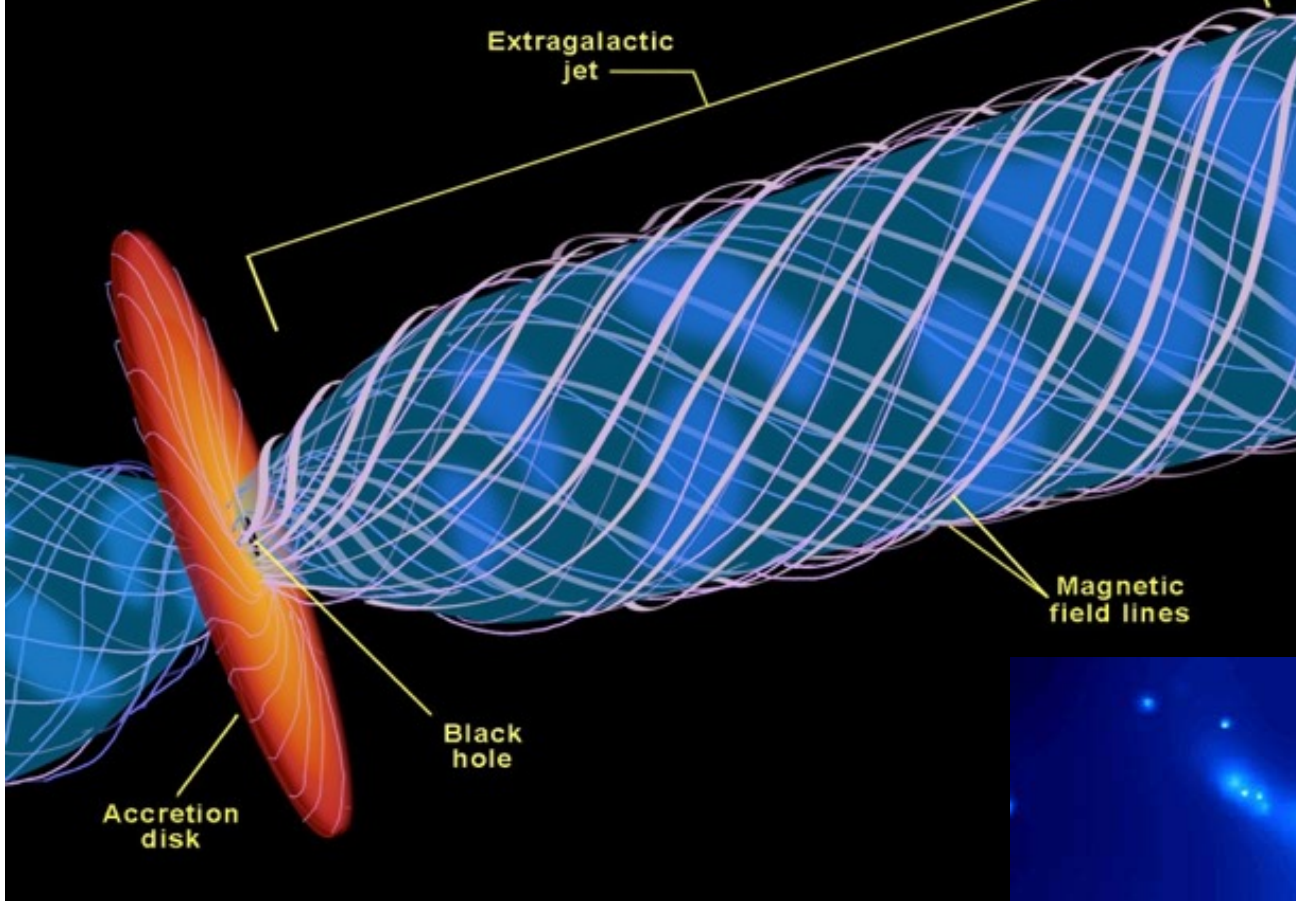
Xu et al. 2015 ApJ Letters 799, L28



SS433 precession jets



Formation of extragalactic jets from black hole accretion disk



A model of Blazar

Fermi's '**Stochastic Acceleration**'
(large synchrotron radiation loss)



Coherent **wakefield acceleration**
(no limitation of the energy)

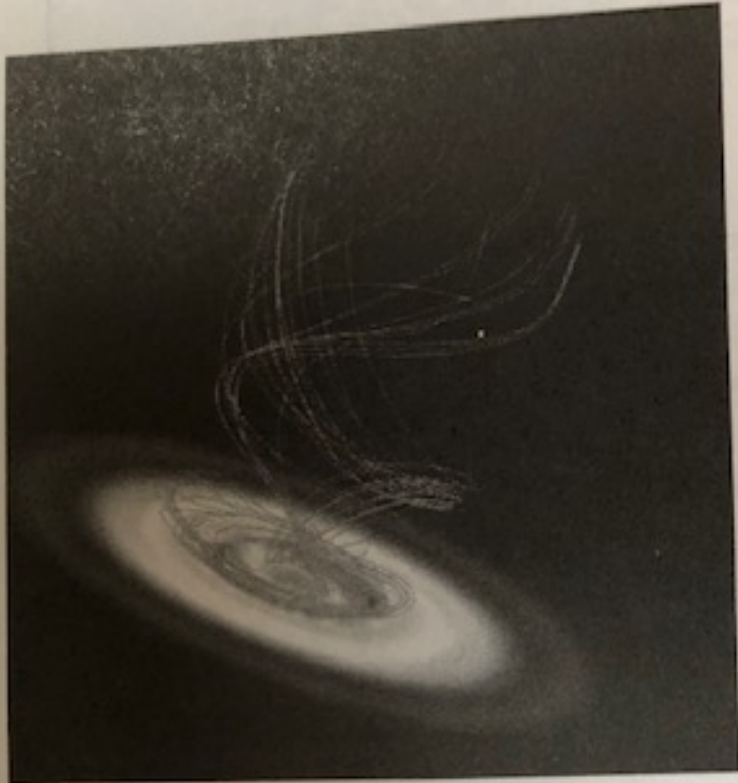
Nature's **LWFA** : Blazar jets

extreme high energy cosmic rays ($\sim 10^{21}$ eV)
episodic γ -ray bursts observed
consistent with **LWFA** theory

Ebisuzaki-Tajima (2014)



Magneto-Rotational Instability (MRI)



Accretion disk
rotating plasma
B-fields

Balbus-Hawley (1991)

FIGURE 4.31 (a) Magnetic field lines and equatorial density; (b) Projection of magnetic field lines (Matsumoto et al. 1991).

Matsumoto Tajima (1995)

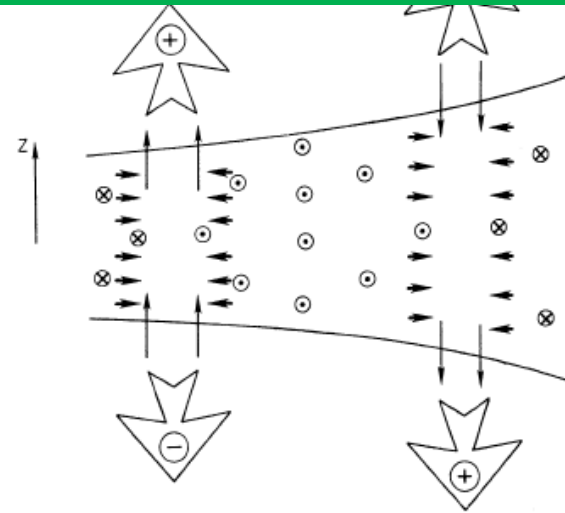
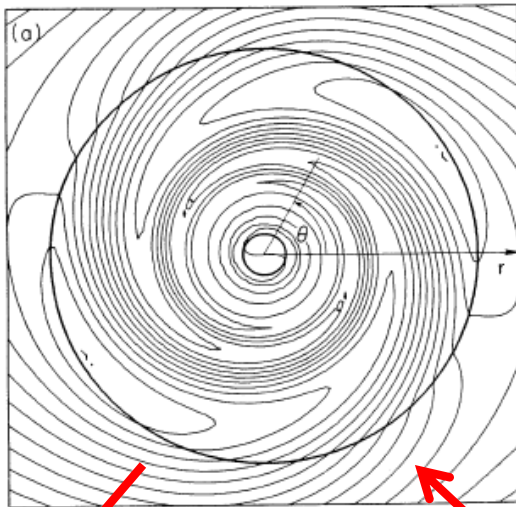
...ing magnetized disks (magnetic Papaloizou-Pringle instability) is observed; (iii) a helical ... rotation spe

Tajima, Shibata (1997)

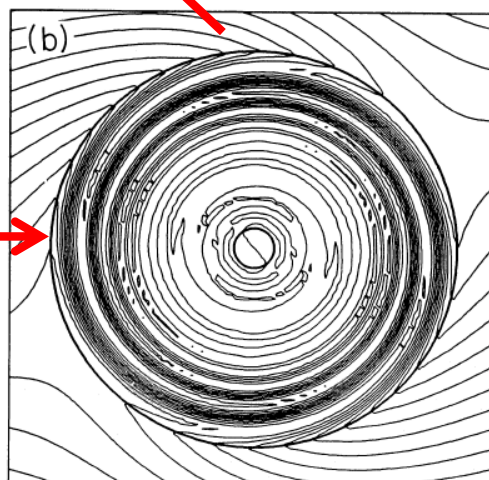
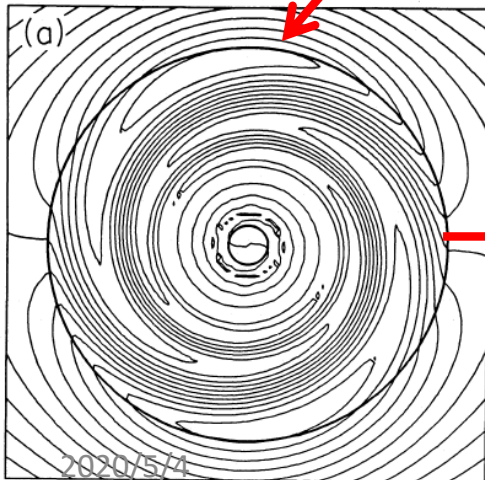
Eruption of magnetic field in an accretion disk

A Burst of Electromagnetic Disturbance

low



growing



high

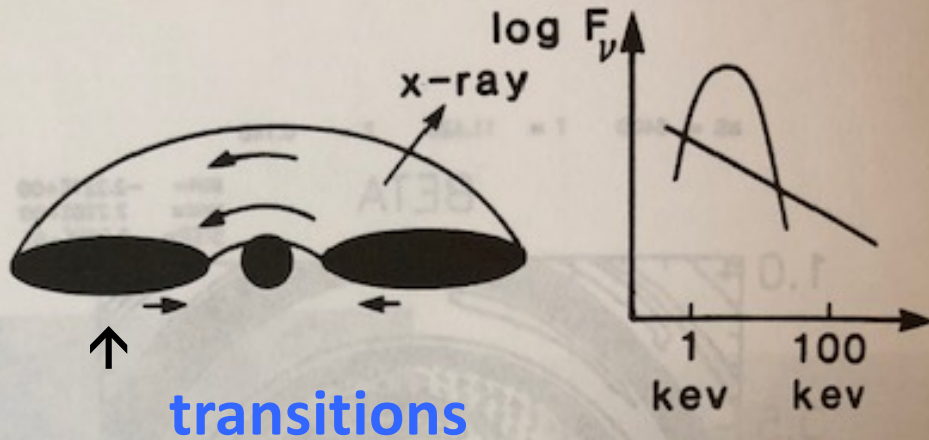
Transition between
high (eruption) and
low (recovery) states

Tajima and Gilden 1987, ApJ 320, 741-745

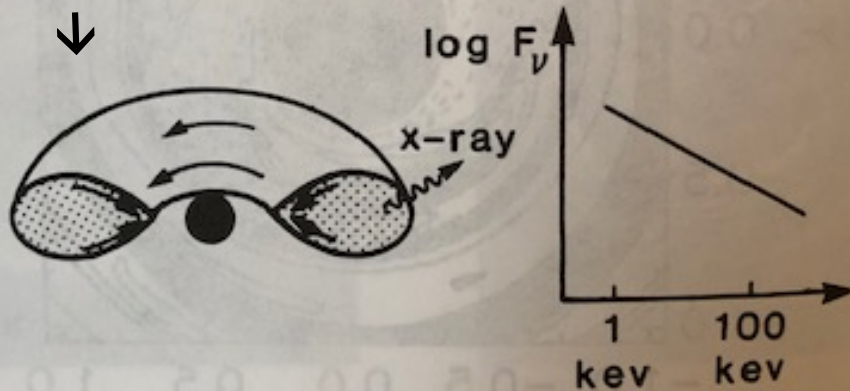
Haswell, Tajima, and Sakai, 1992, ApJ, 401, 495-507

Episodic transitions of accretion disks with B

● High State (Soft State)



● Low State (Hard State)



(Quasi-)steady state accretion disks

cf.

Episodic accretion disks
“breathing” disks

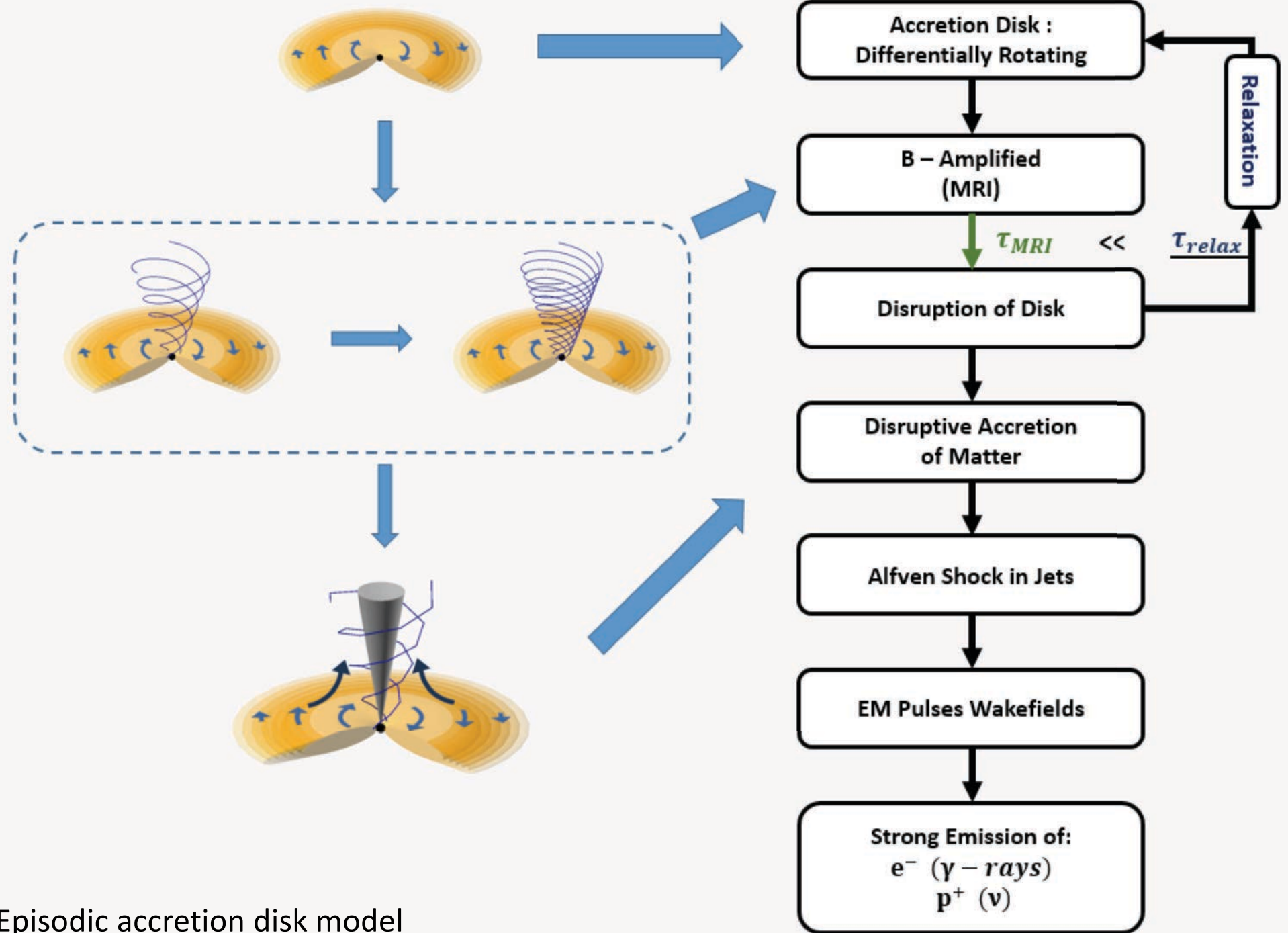
transitions between the states
soft state ($\beta \gg 1$)



hard state ($\beta \sim 1$)

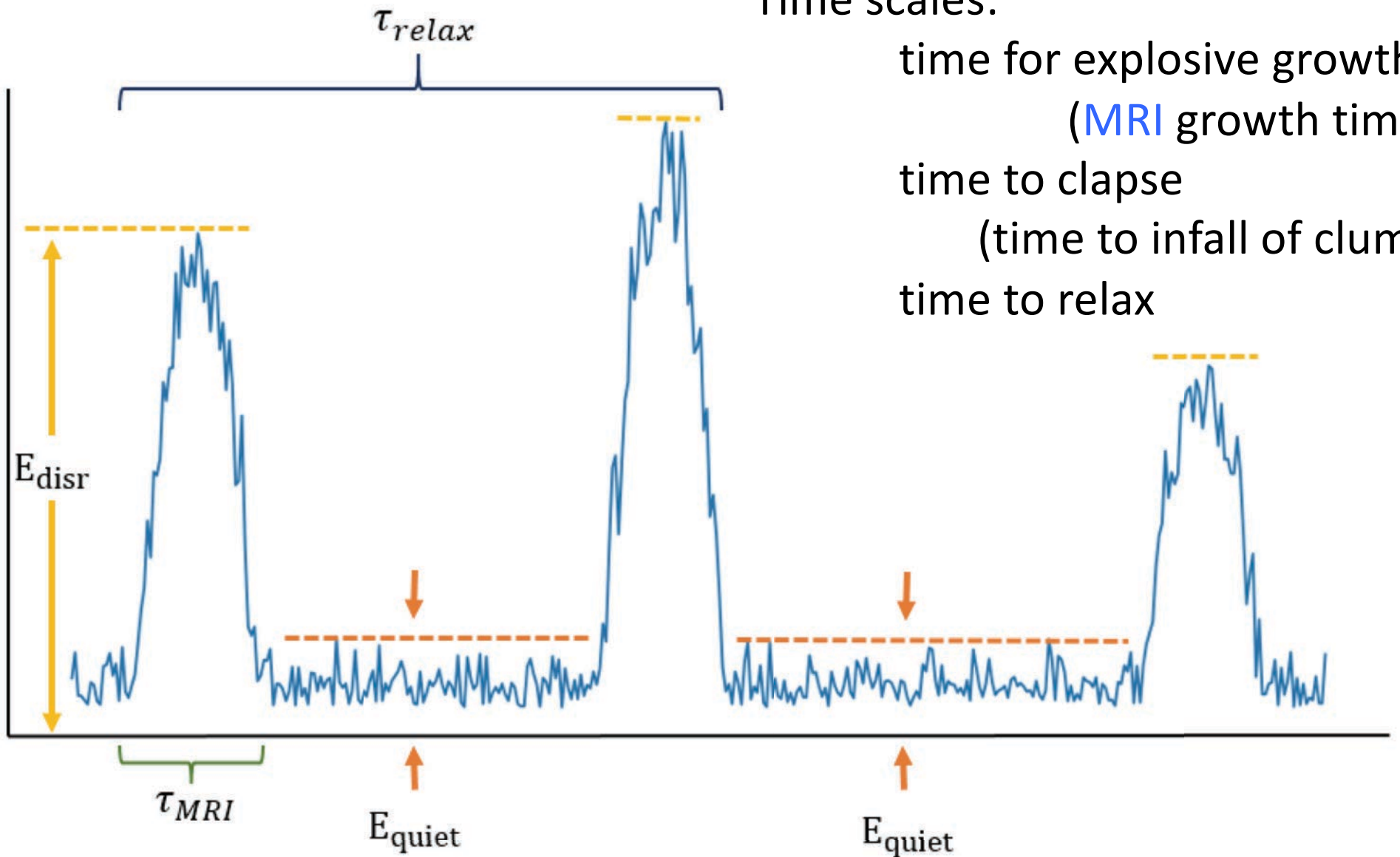
Two states of accretion disks: “High state” (soft state) vs. “Low state” (hard state): the spectra

It is known that accretion disks in black hole candidates have two spectral states (Miyamoto et al., 1984). One is the high state and the other is the low state. In the high state, the spectrum has blackbody component which can be explained by emission from the inner accretion disks. On the other hand, in the low (or hard) state, the spectrum has power-law component which may come from optically thin accretion disk (Fig. 4.33). Other



Episodic accretion disk model
 Abazajian, Tajima, Ebisuzaki

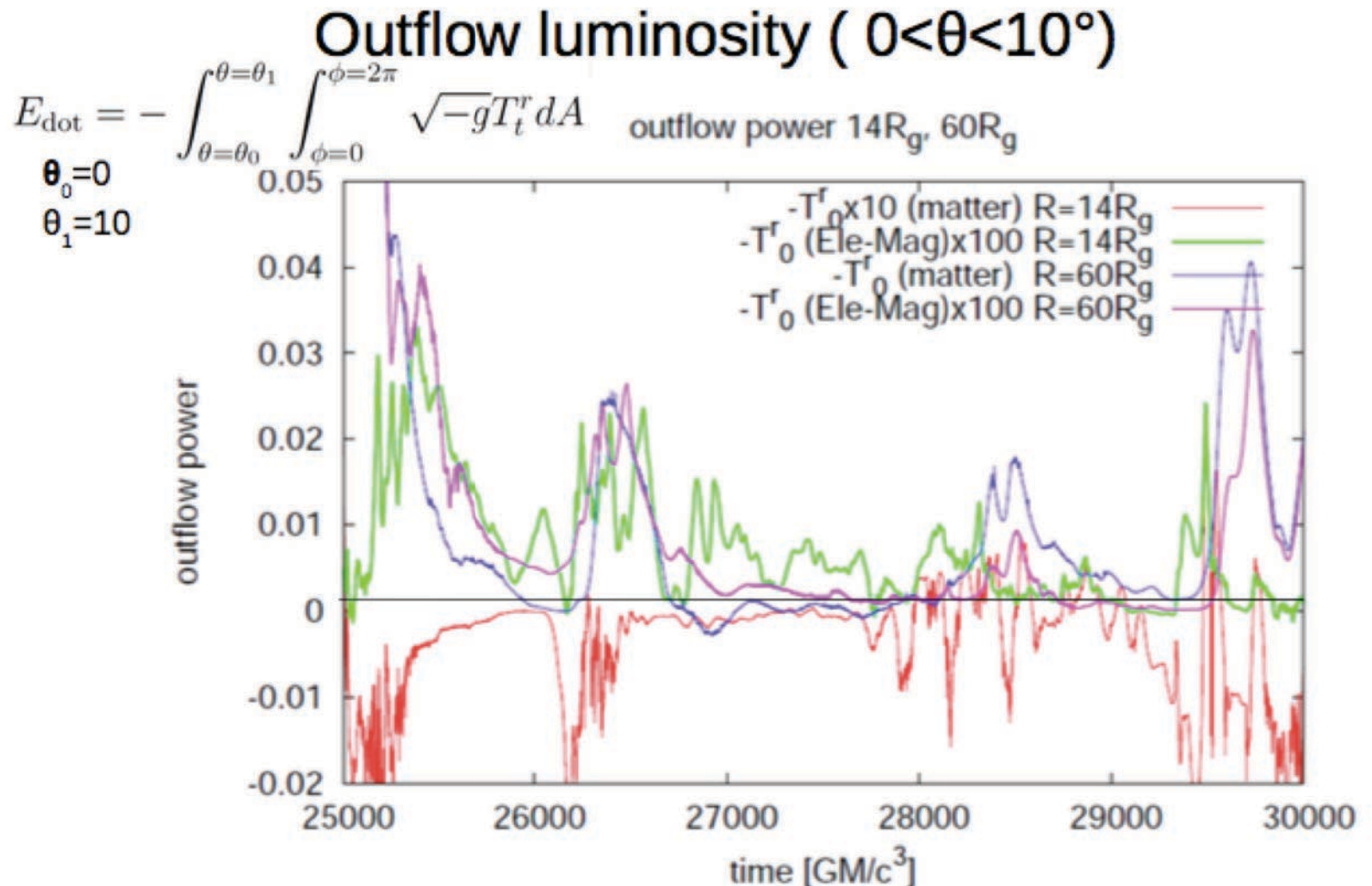
Episodic eruption of accretion disk



Time scales:

- time for explosive growth
(MRI growth time)
- time to clapse
(time to infall of clump)
- time to relax

General Relativistic MHD simulation of accretion disk + jets: episodic feature



Short time variability ($\Delta t \sim$ a few tens GM/c^3) in electromagnetic components (green and pink) : Good agreement with Ebisuzaki & Tajima(2014) $t_{var} \sim M$
 \Rightarrow possible origin for flares in blazars,
 strong Alfvén wave mode \Rightarrow Application to wake field acc. for UHECRs

Intense Alfvén Shock from root of jet

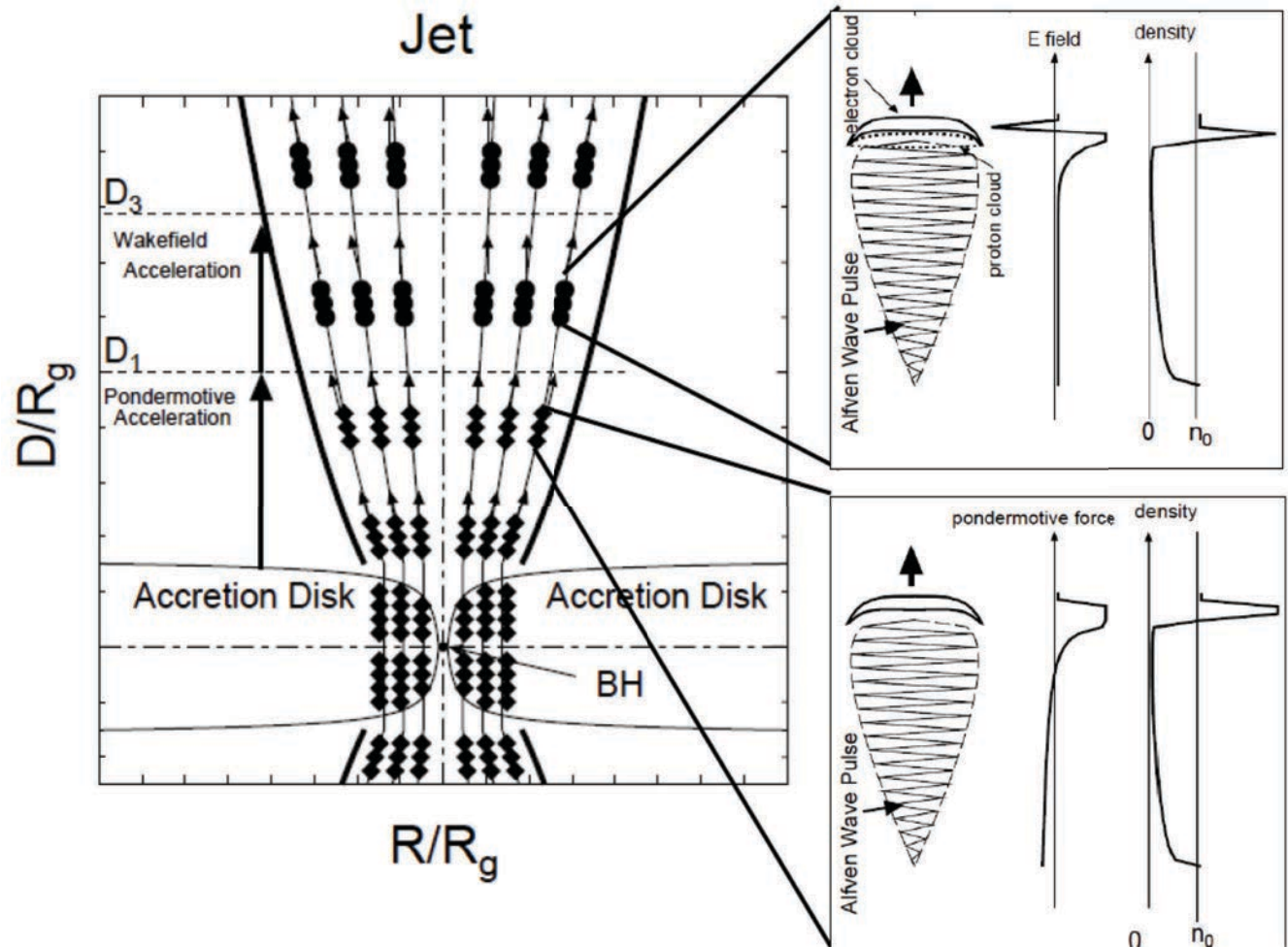
→ Intense **EM** pulse

→ **wakefield** generation → **Electron** acceleration

→ **γ burst** ion

UHECR

← n_e decreases



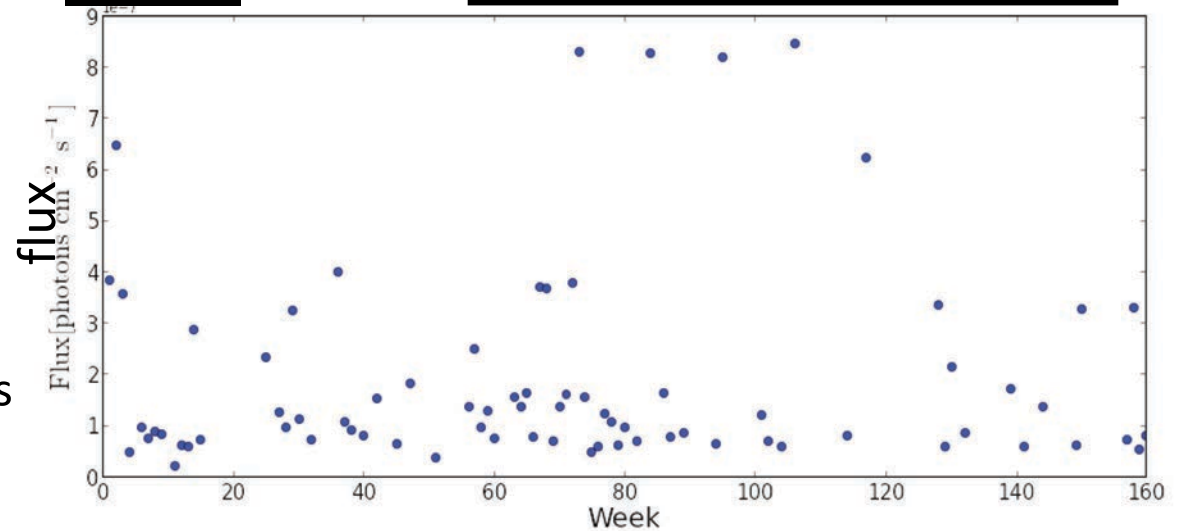
Blazar shows anti-correlation between γ burst flux and spectral index

Blazar: A00235+164

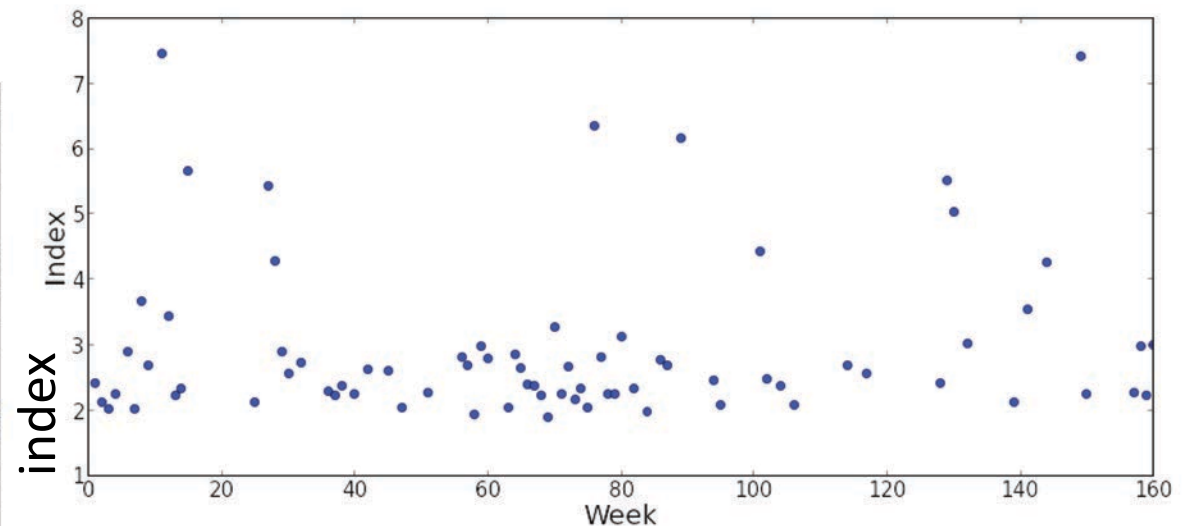
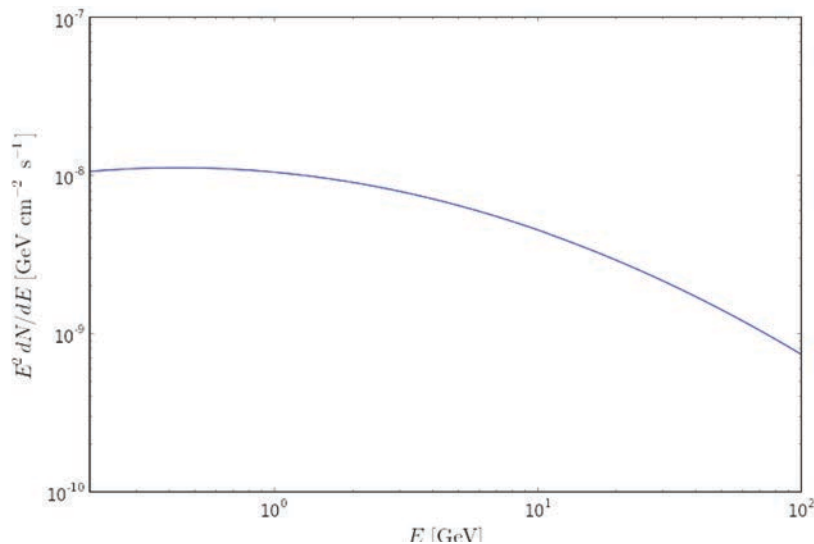
$M \sim 10^8 M_{\text{Sun}}$

Rise time < week (less than a unit),
Period between bursts $\sim > 10$ weeks
Spectral index $\Rightarrow 2$

(\sim Ebisuzaki/Tajima theory)



\rightarrow all quantitatively consistent with Wakefield theory



time

N. Canac, K. Abazajian (2020)

Again, Anti-correlation even in a bigger blazar

Blazar: 3C454.3

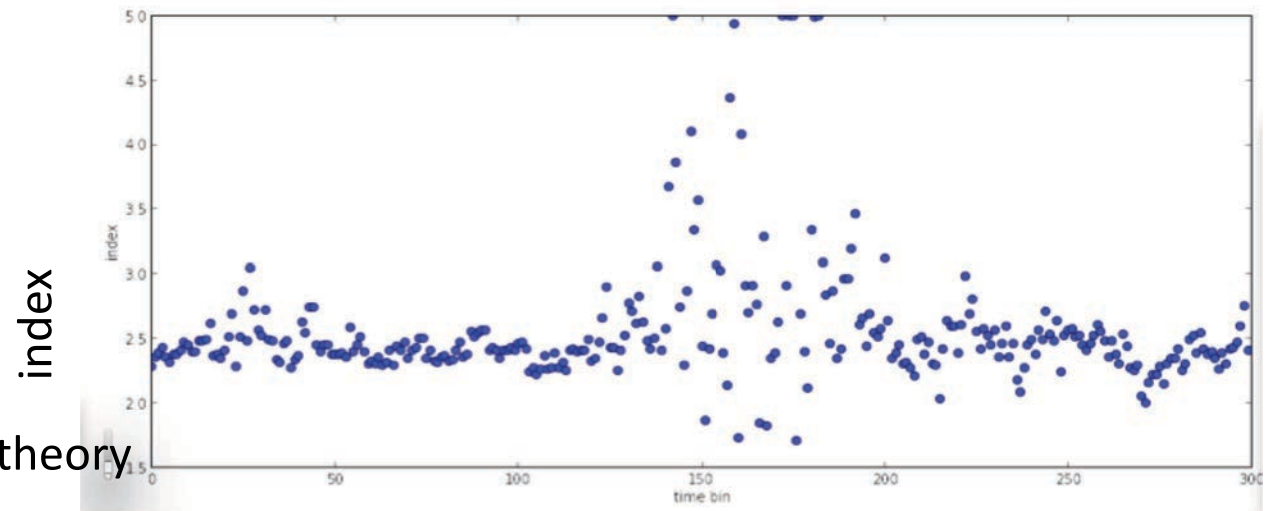
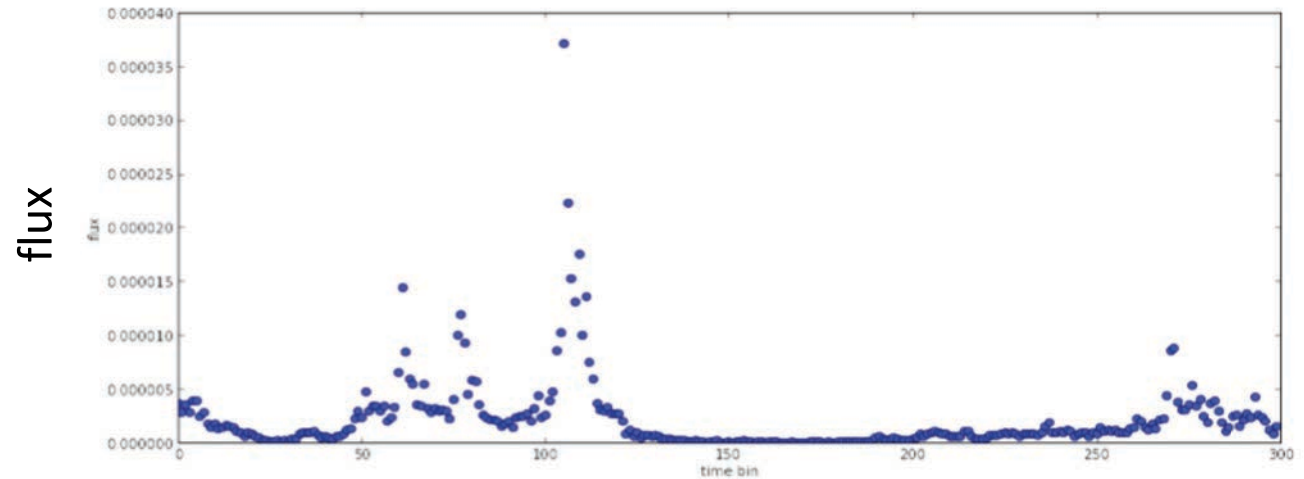
$M \sim 10^9 M_{\text{Sun}}$

Same anti-correlation as
AO0235+164

The rise time and burst periods
a lot longer (by an order of
magnitude)

Quantitative agreement and
correct scaling with Blazar mass
with (broader sense of) Wakefield theory
(Ebisuzaki/Tajima)

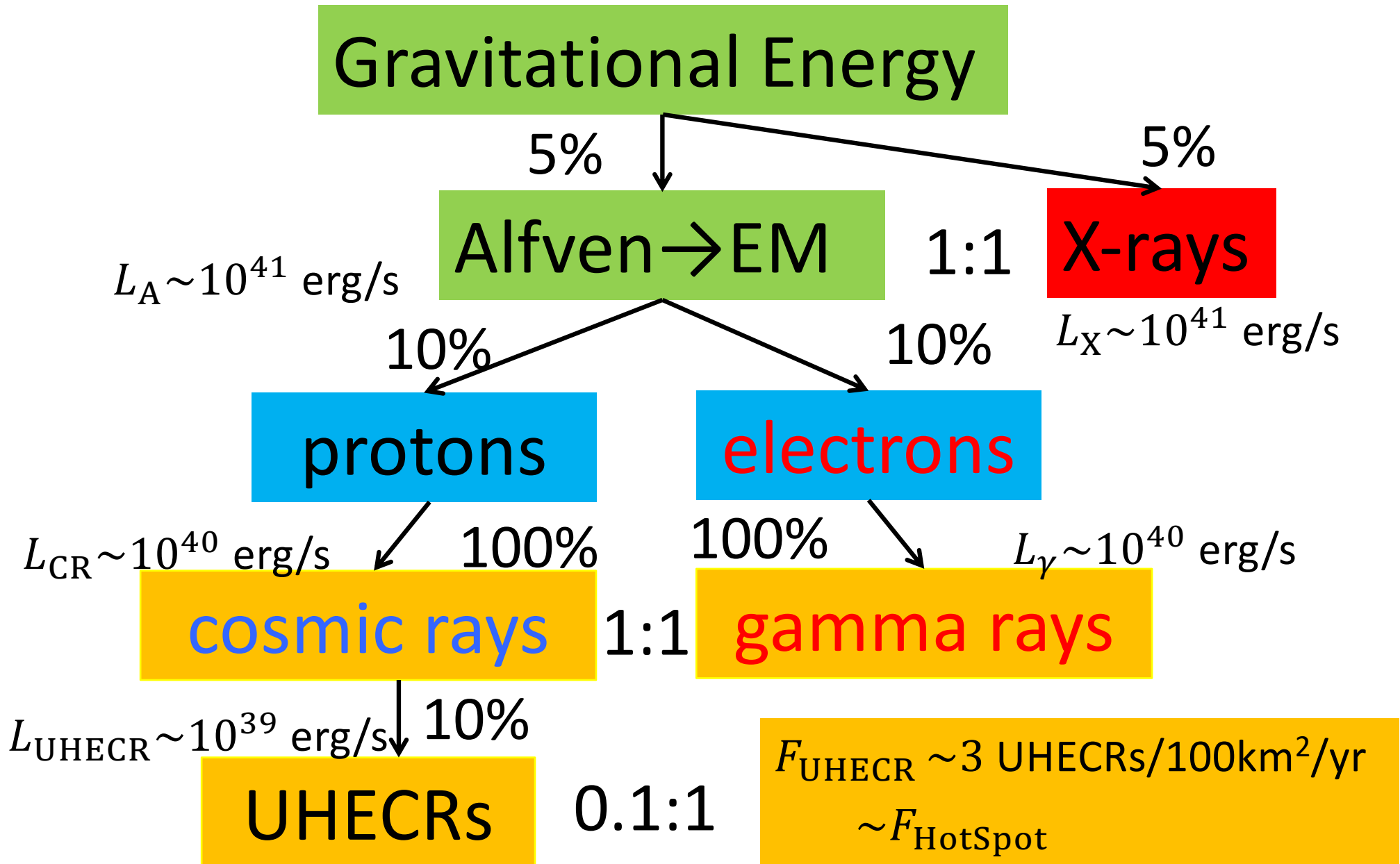
period $\sim M$; luminosity $\sim M$



time

N. Canac, K. Abazajian (2020)

Energy release by **wakefield** (e.g. M82 X-1)





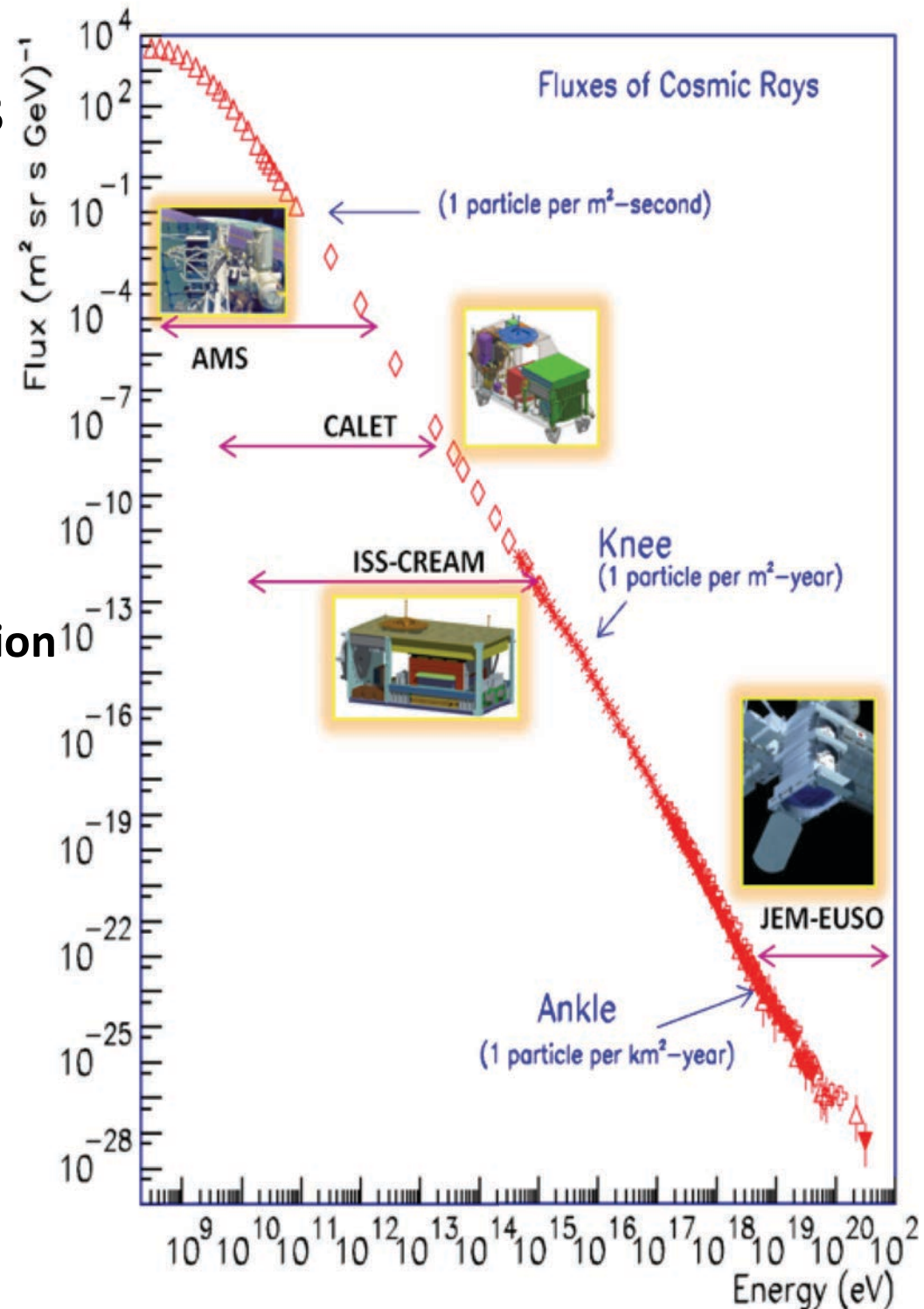
Nature's wakefield accelerators in cosmos

3. Pinpointed high energy cosmic rays (and neutrinos)

Ultrahigh Energy Cosmic Rays (UHECR)

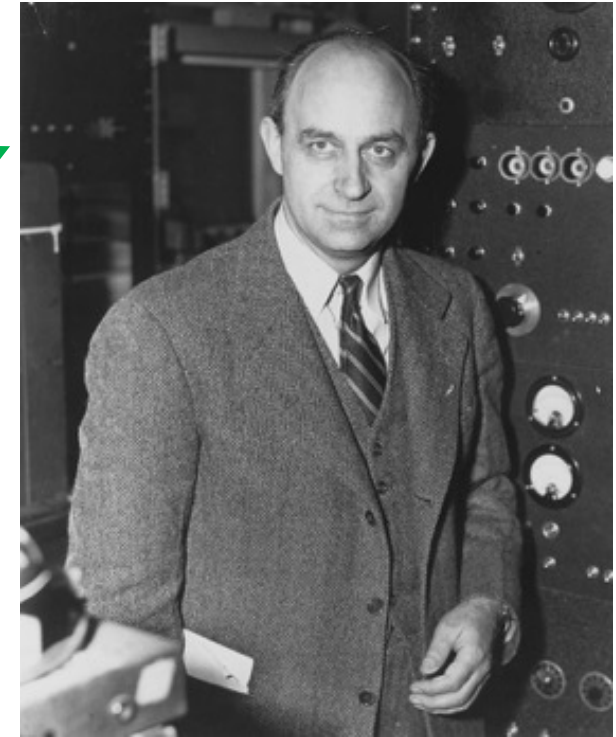
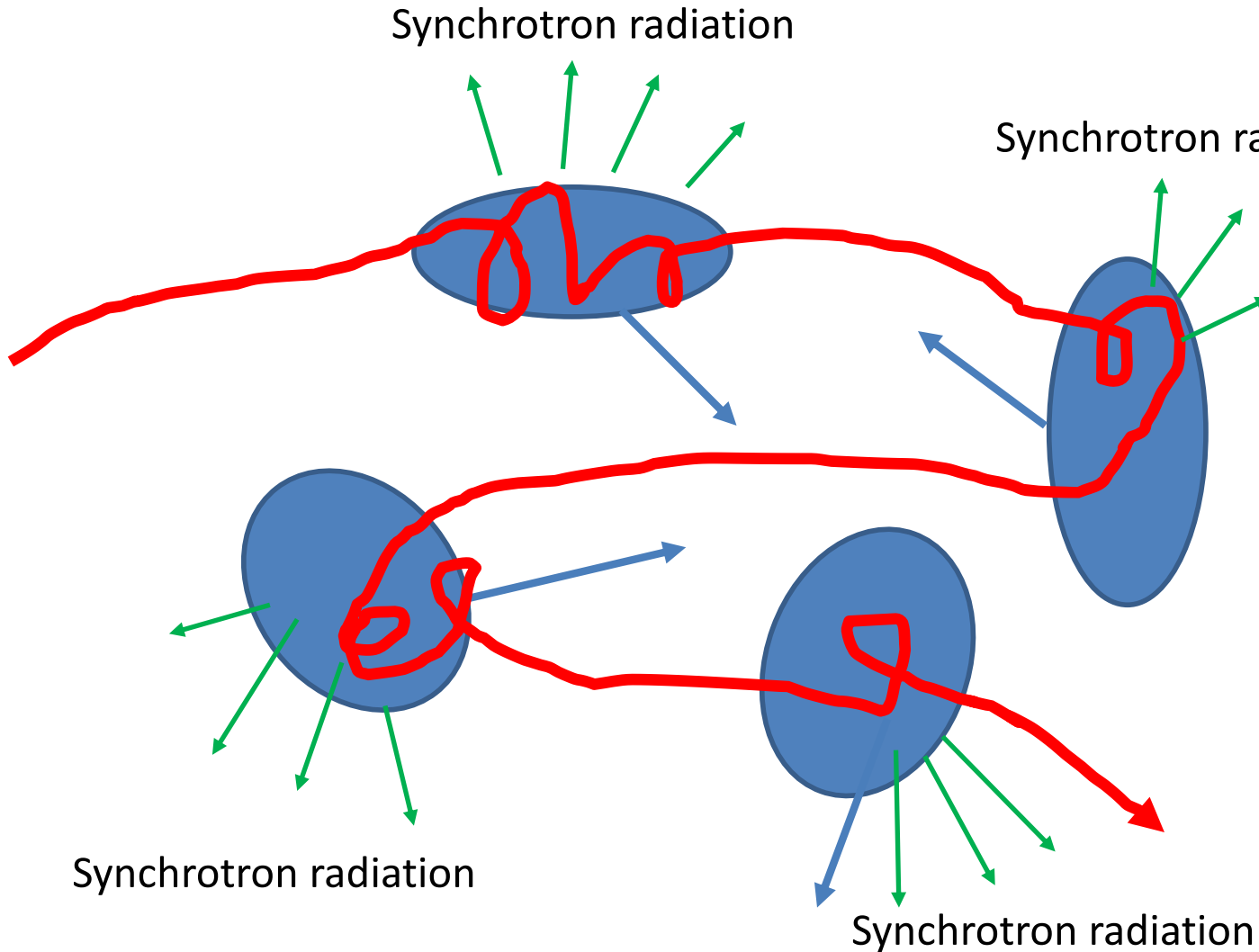
Fermi mechanism runs out of steam
beyond 10^{19} eV
due to synchrotron radiation

Wakefield acceleration
comes in rescue
prompt, intense, linear acceleration
small synchrotron radiation
radiation damping effects?



Fermi mechanism

incoherent
requires bending → synchrotron loss



Enrico Fermi

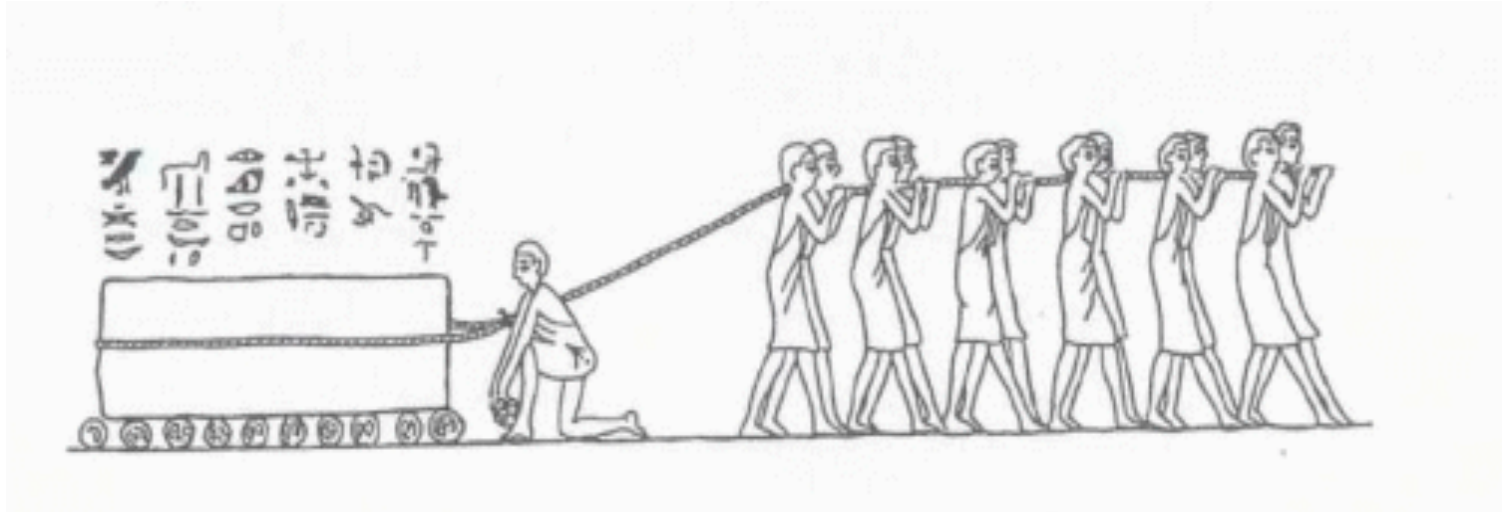
By Department of Energy. Office of Public Affairs

E. Fermi, ApJ 119 (1954) 1.

Plasma's **Collective** Force / Modes

Collective force $\sim N^2$ (nonlinear \leftarrow linear force $\sim N$)

Coherent and smooth structure (not stochastic)



enhancement by $10^3 - 10^4$ (even by 10^{6-12}) \gg interaction of one particle x one particle

Collective mode delivery (EM x plasma x B) \leftrightarrow **long-ranged** force (gravity, EM)

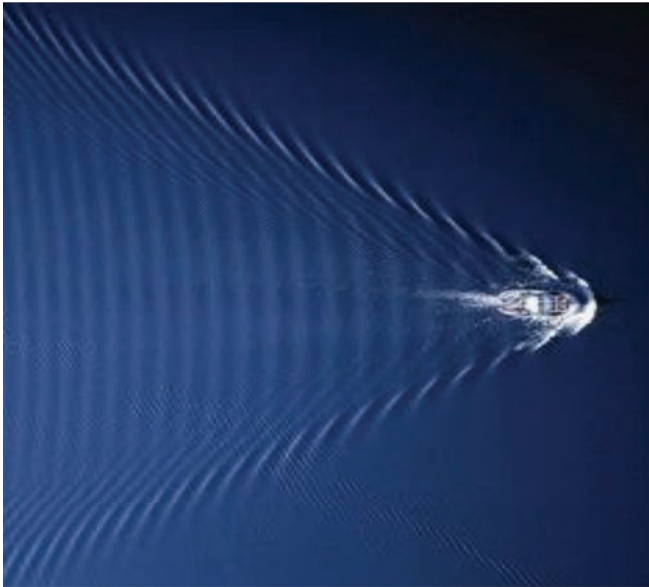
what difference?

e.g. **jet**

e.g. galaxy-galaxy interaction

Laser Wakefield (LWFA):

Wake phase velocity \gg water movement speed
maintains **coherent** and **smooth** structure



Tsunami phase velocity becomes ~ 0 ,
causes **wavebreak** and **turbulence**

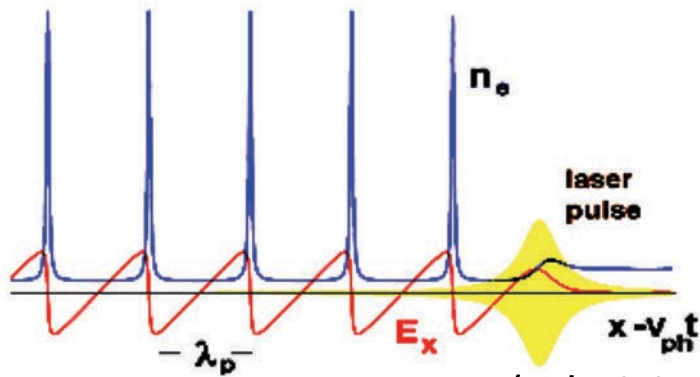


VS

Strong beam (of laser / particles) drives plasma waves to saturation amplitude: $E = m\omega v_{ph} / e$

No wave breaks and wake **peaks** at $v \approx c$

Wave **breaks** at $v < c$



← relativity
regularizes
(*relativistic coherence*)

$a_0 \sim 10 \gg 1$ (relativistic wave)



Relativistic coherence enhances beyond the Tajima-Dawson field $E = m\omega_p c / e$ (\sim GeV/cm)

High phase velocity paradigm

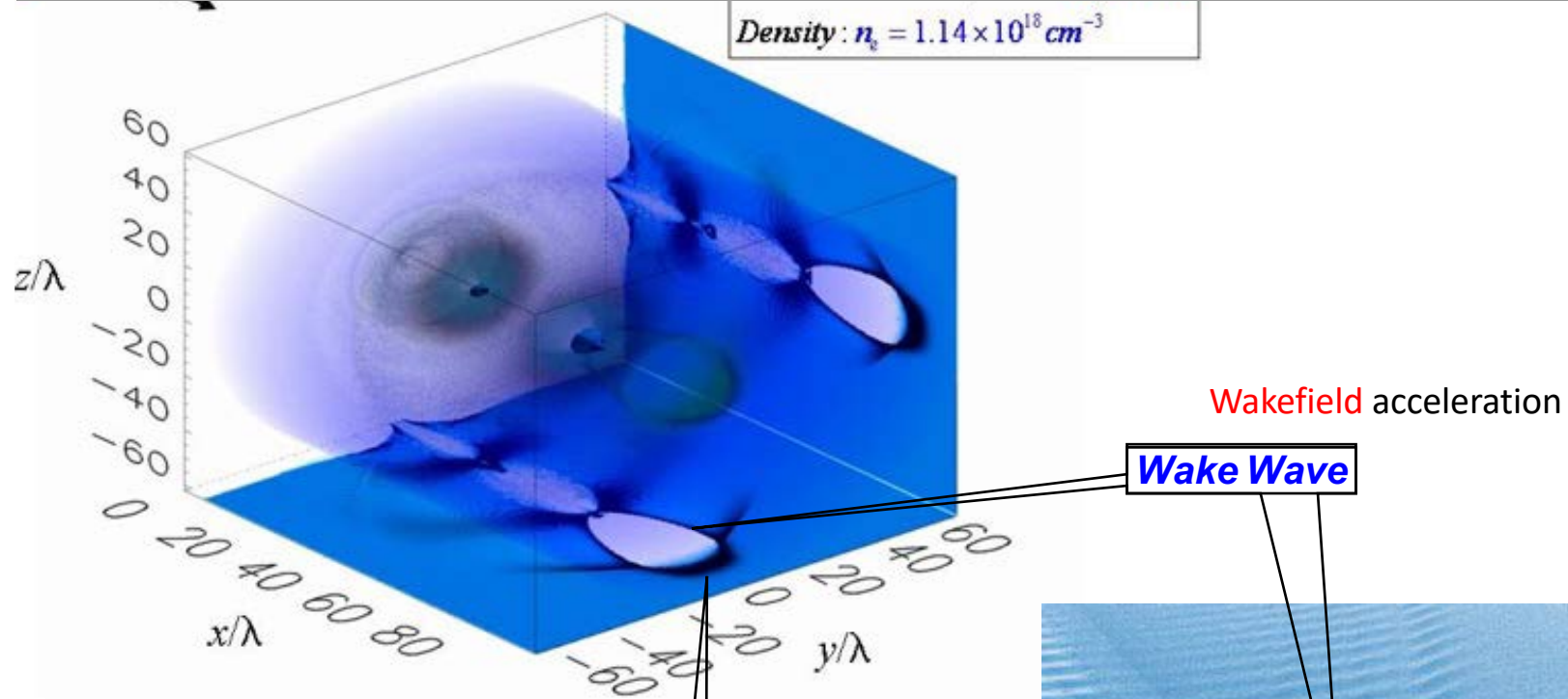
unstable, chaotic

robust, coherent

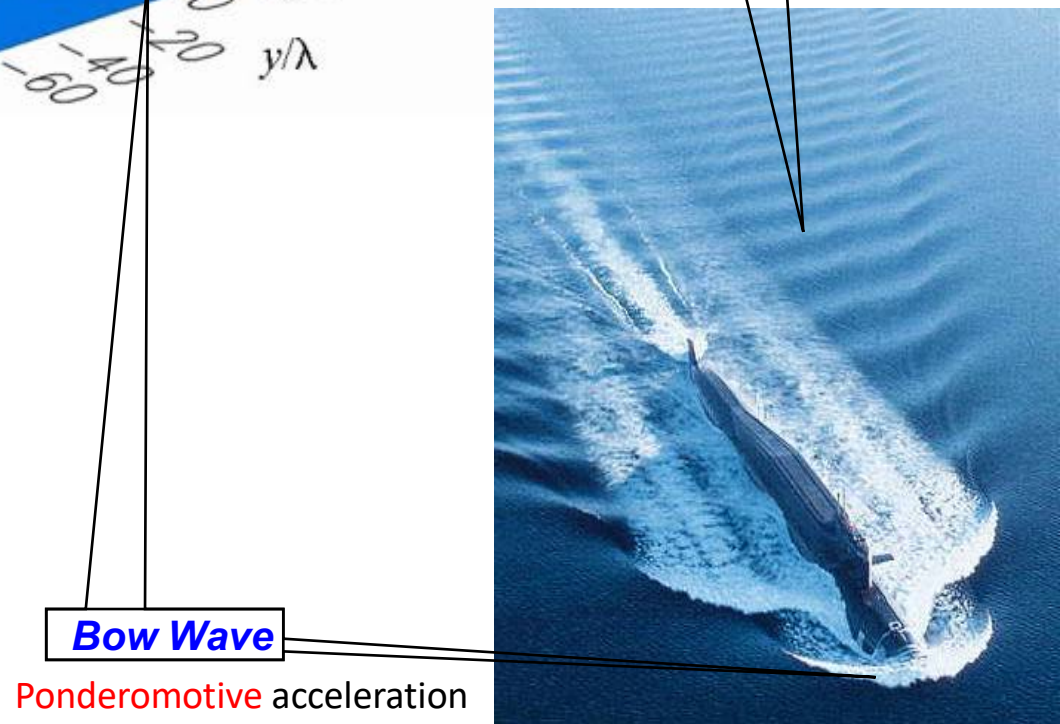
Low phase velocity	High phase velocity
Plasma tends to be unstable	Stable state exists (Landau-Ginzburg state)
$v_{ph} \sim v_{th}$	$v_{ph} \gg v_{th}$
Mode interacts with bulk plasma (Landau resonance)	Mode insulated from bulk plasma
Mode-mode coupling <ul style="list-style-type: none"> ➔ More modes ➔ More turbulence 	Mode maintains coherence
Strongly nonlinear regime (large Reynolds' number) ➔ strong turbulence	Strongly nonlinear regime ➔ strongly coherent <u>Relativistic effects</u> further strengthen coherence
<u>Plasma fragile</u> ➔ anomalous transport, structure disintegration	Plasma cannot be destroyed, structures are formed. Violence tolerated
Trapping: $v_{tr} \lesssim v_{th} \sim v_{ph}$ $x_{tr} = \sqrt{\frac{cE}{B} \frac{L_s}{k_y v_{\parallel}}} \quad 22$	Trapping: $v_{tr} = \sqrt{qE/mk} \quad 13$ <p>If wave pumped, v_{tr} increases until $v_{tr} \sim v_{ph} \gg v_{th}$ ➔ acceleration or injection</p> <p>Tajima-Dawson saturation: $E_{TD} = \frac{m\omega_p c}{e}$ </p>
Characteristic structure: Sheath	Characteristic structure: <u>Wake</u>
Energy gain: by coherent accumulation of electron charges of the sheath (energy amplification of sheath charge accumulation $2\alpha + 1$ (coherence parameter α) ¹⁸	Energy gain: by energy amplification over the trapping width $v_{tr} \sim v_{ph}$ (Lorentz transform factor $2\gamma^2 = 2 n_{cr}/n_e$)

Tajima, Yan, Ebisuzaki (2020)

Laser-driven Bow and Wake

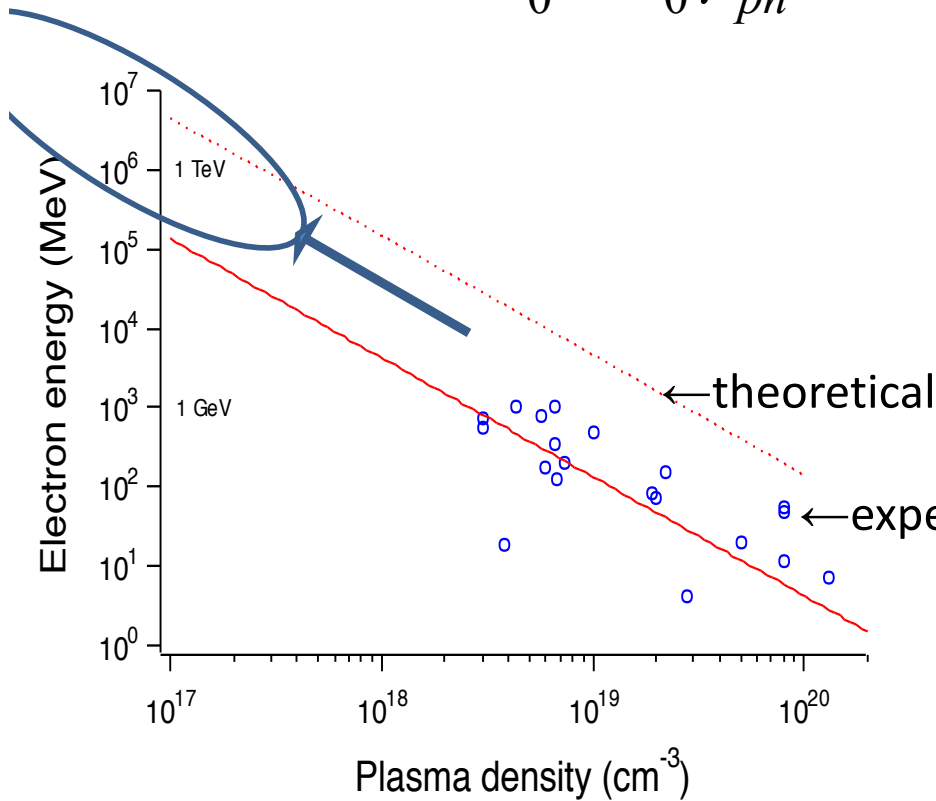


(Bulanov, Esirkepov)



Universal Theory of Wakefield toward extreme energy

$$\Delta E \approx 2m_0c^2 a_0^2 \gamma_{ph}^2 = 2m_0c^2 a_0^2 \left(\frac{n_{cr}}{n_e} \right), \quad (\text{when 1D theory applies})$$



In order to avoid wavebreak,

$$a_0 < \gamma_{ph}^{1/2},$$

where

$$\gamma_{ph} = (n_{cr} / n_e)^{1/2}$$

$$n_{cr} = 10^{21} \text{ (fs photon (laser))}$$

$$= 10 \text{ (} 10^3 \text{ s wave in disk)}$$

$$n_e = 10^{18} \text{ (gas)}$$

$$= 10^{-2} \text{ (gas in the jet)}$$

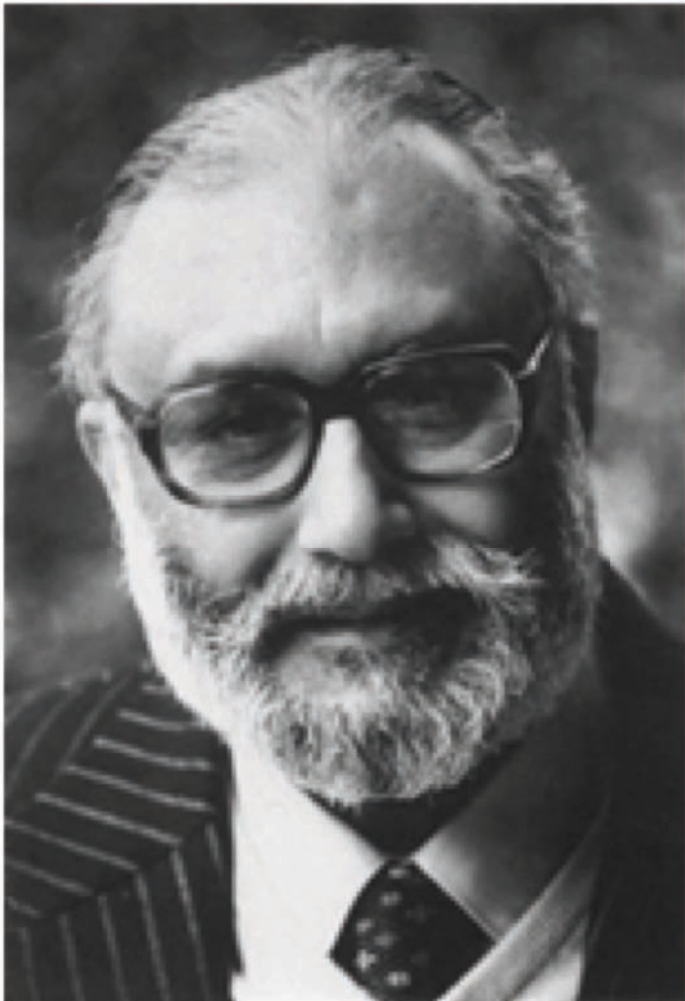
$$L_d = \frac{2}{\pi} \lambda_p a_0^2 \left(\frac{n_{cr}}{n_e} \right), \quad L_p = \frac{1}{3\pi} \lambda_p a_0 \left(\frac{n_{cr}}{n_e} \right),$$

dephasing length

pump depletion length

The late Prof. **Abdus Salam** (1981)

At ICTP Summer School (Trieste, 1981), Prof. Abdus Salam summoned me and discussed about [laser wakefield](#) acceleration.



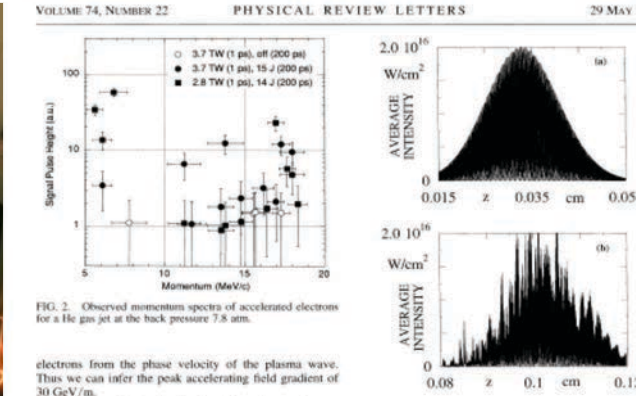
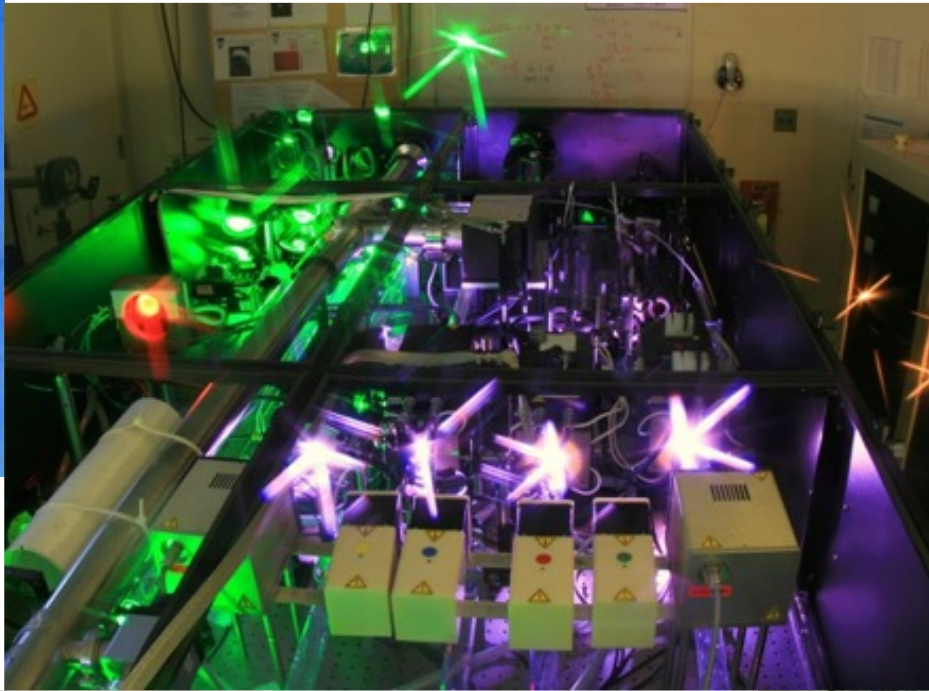
Salam: *‘Scientists like me began feeling that we had less means to test our theory. However, with your laser acceleration, I am encouraged’.* **(1981)**

He organized the Oxford Workshop on [laser wakefield](#) accelerator in 1982.

Demonstration (1994), realization, and applications of laser wakefield accelerators

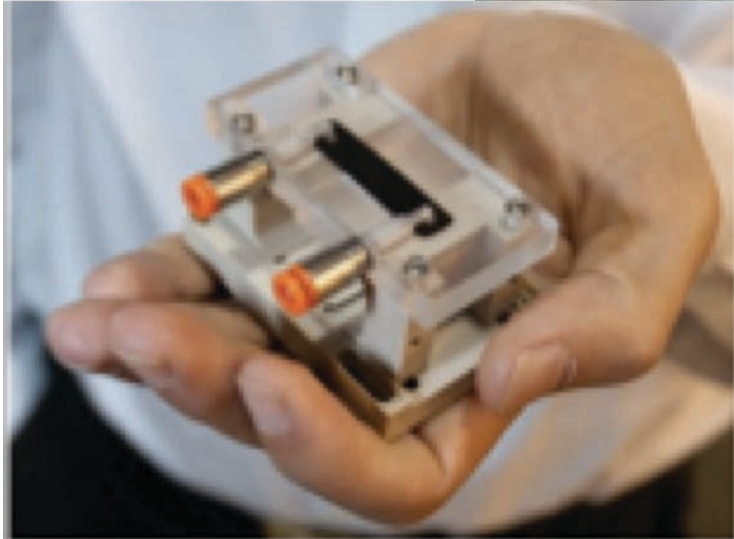


(2004)



Nakajima, et al (1994, 1995)

(Michigan)



4 GeV laser accelerator LBL



3GeV Synchrotron SOLEIL



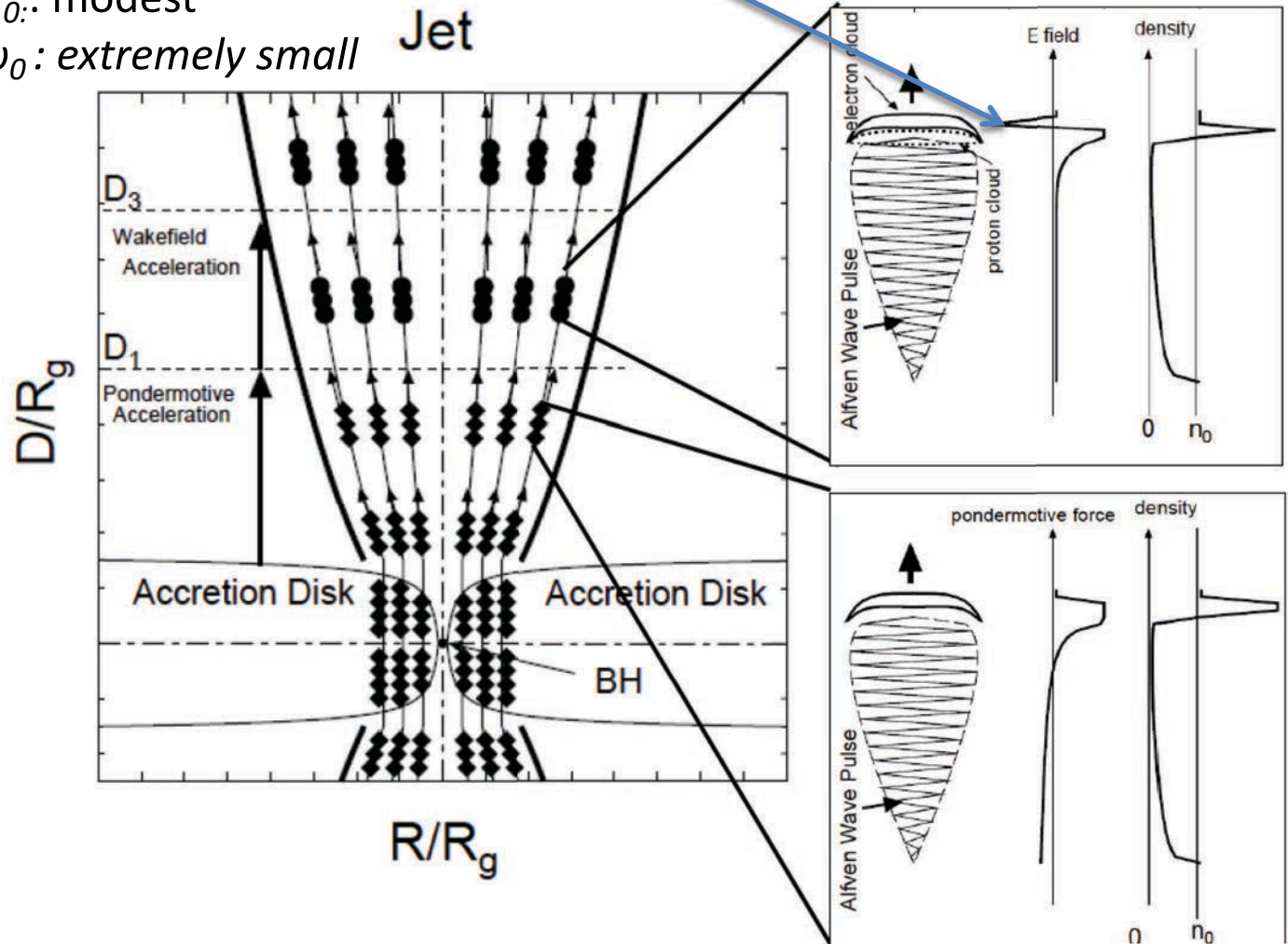
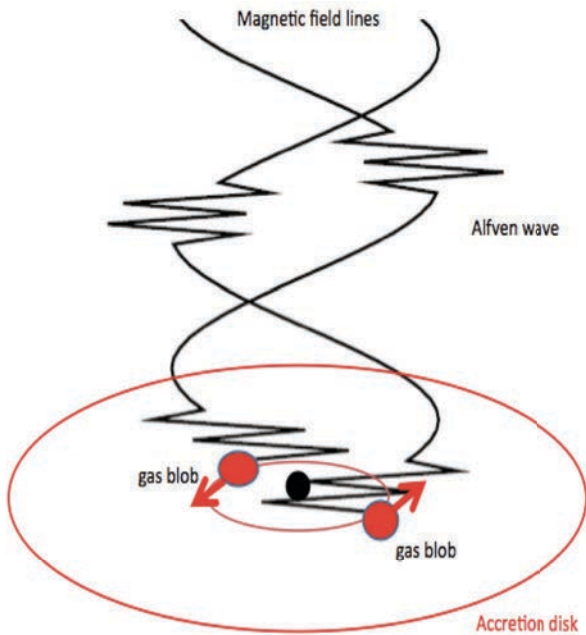
Astrophysical wakefield acceleration:

Superintense Alfvén Shock in the Blackhole Accretion Disk toward ZeV Cosmic Rays ($a_0 \sim 10^6 - 10^{10}$, large spatial scale)

$$a_0 = eE_0 / mc\omega_0 \gg 1$$

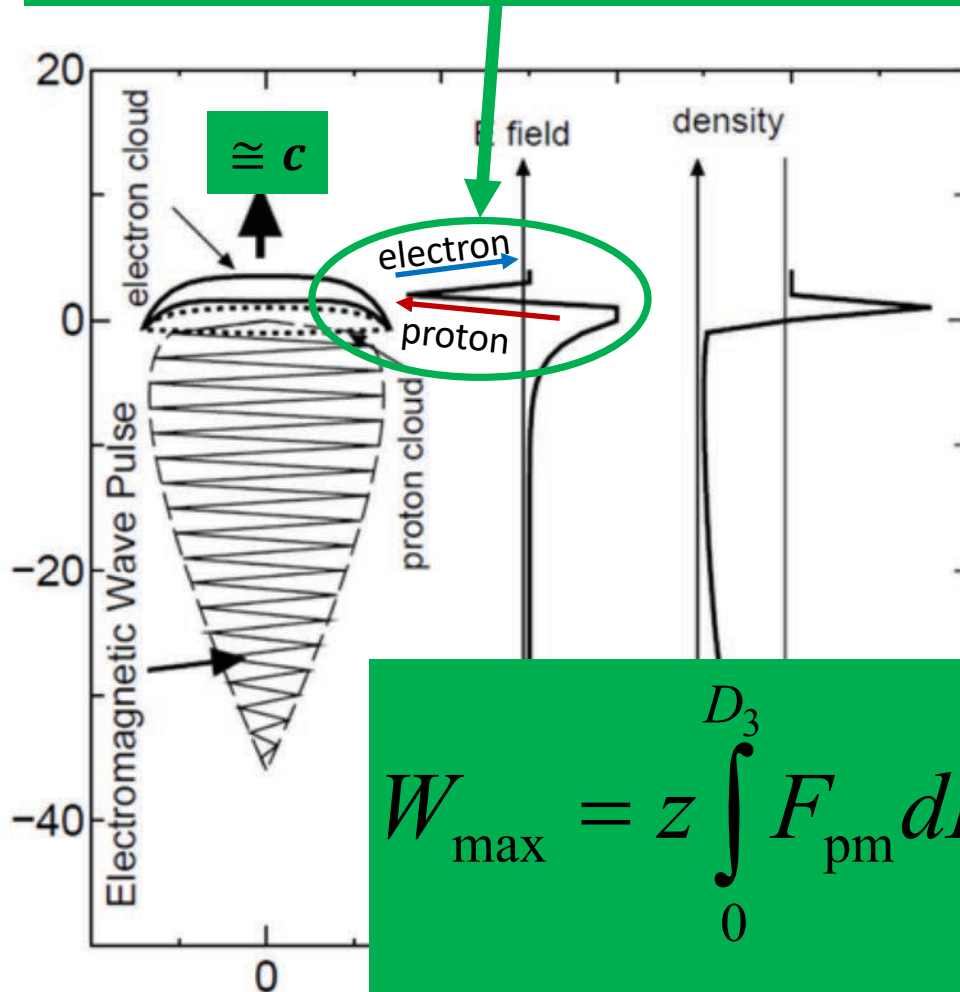
E_0 : modest

ω_0 : extremely small



Wakefield Acceleration

Co-linear acceleration by electrostatic field

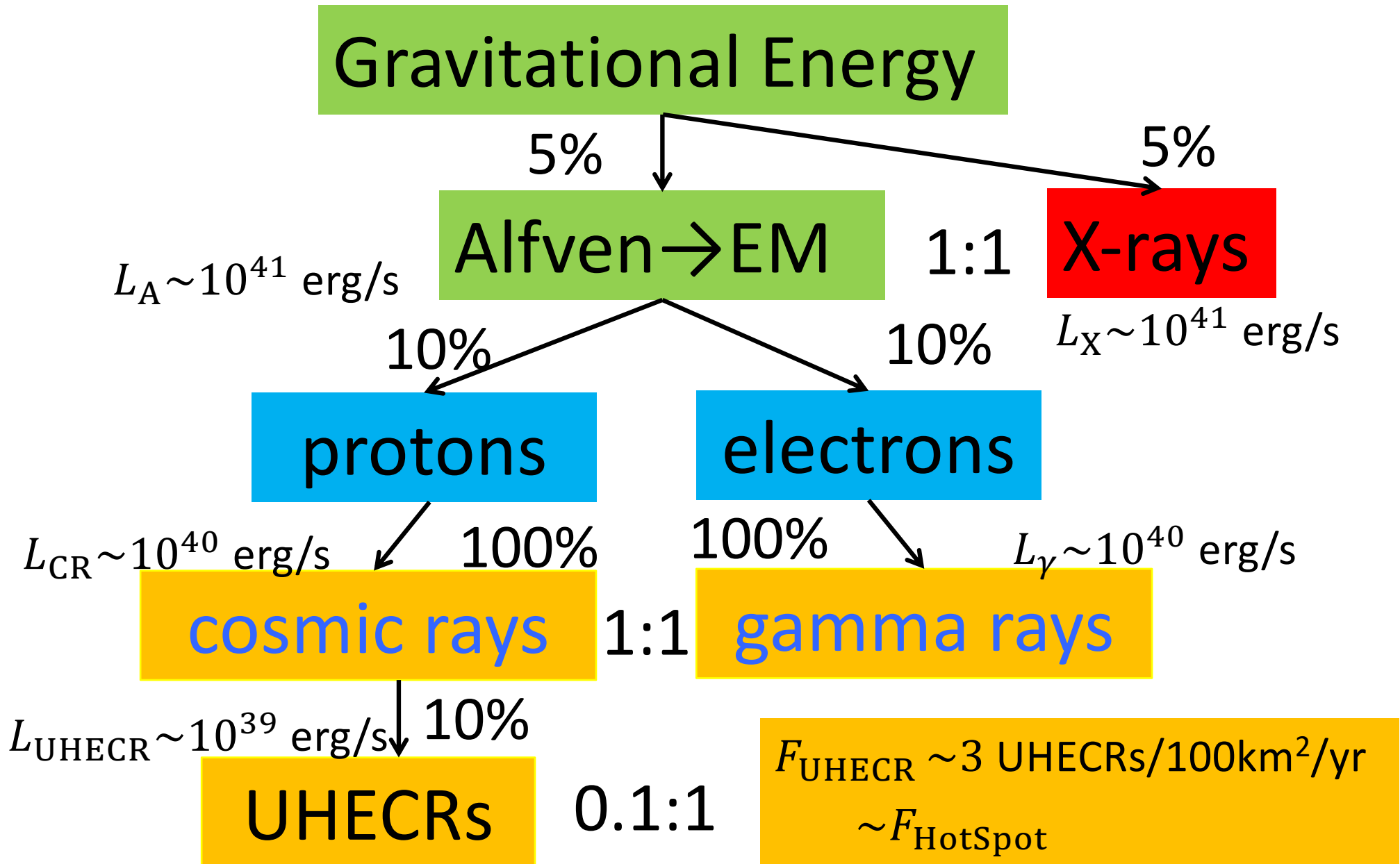


$$W_{\max} = z \int_0^{D_3} F_{\text{pm}} dD$$

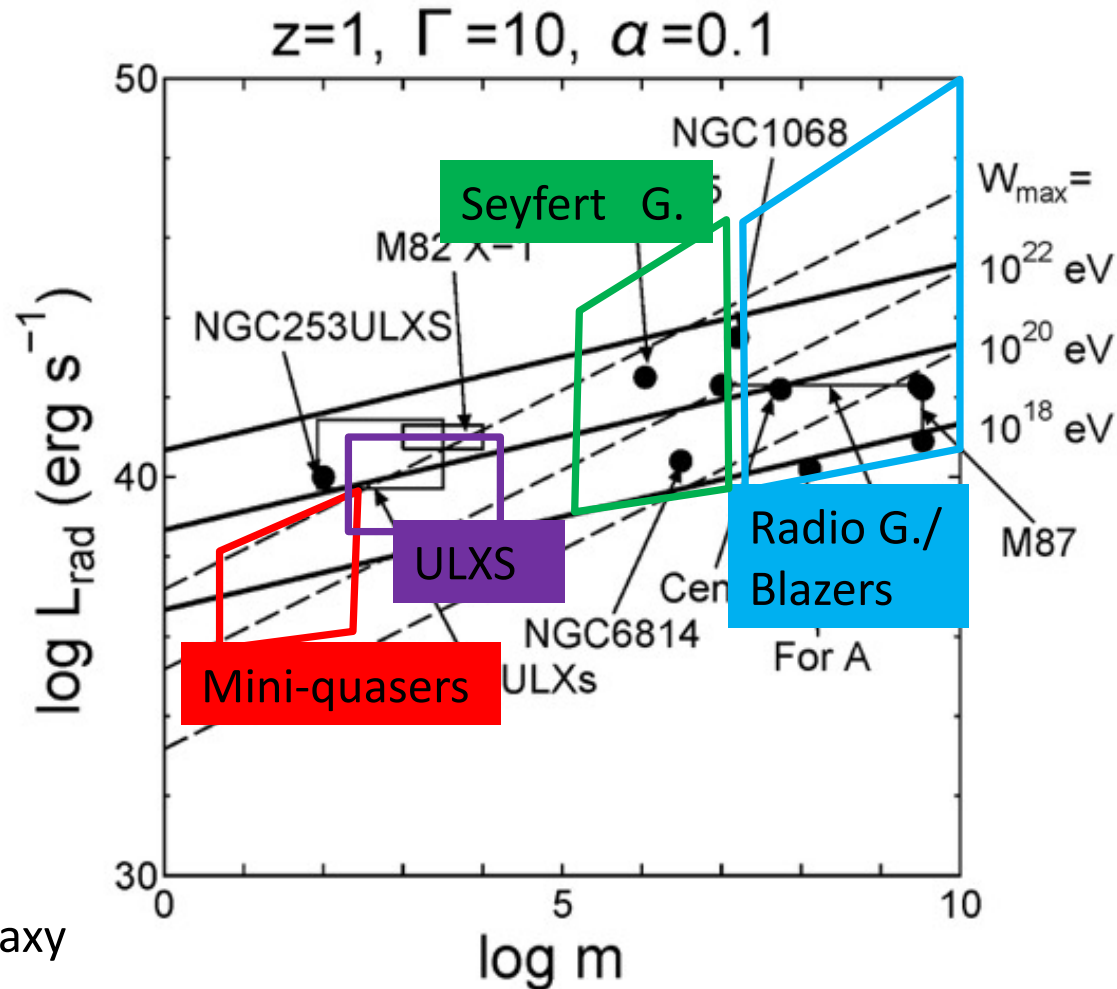
$$F_{\text{pm}} = \Gamma m_e c a_0 \omega_A$$

- **Stable acceleration structure**
 - Coherent and Strong Field
 - Moving in $\cong c$
 - Colinear acceleration
 - across a long length
 - Built in deep in the theory
- **All the messenger channels**
 - Electrons \rightarrow photons (HE, radio)
 - Protons \rightarrow CRs \rightarrow neutrinos
 - Gravitational waves (NS mergers)
- **Variabilities**
 - Caused by disk instability
 - In all messenger channels
 - Violent and simultaneous

Energy release by wakefield



cosmic ray acceleration and gamma-ray emission

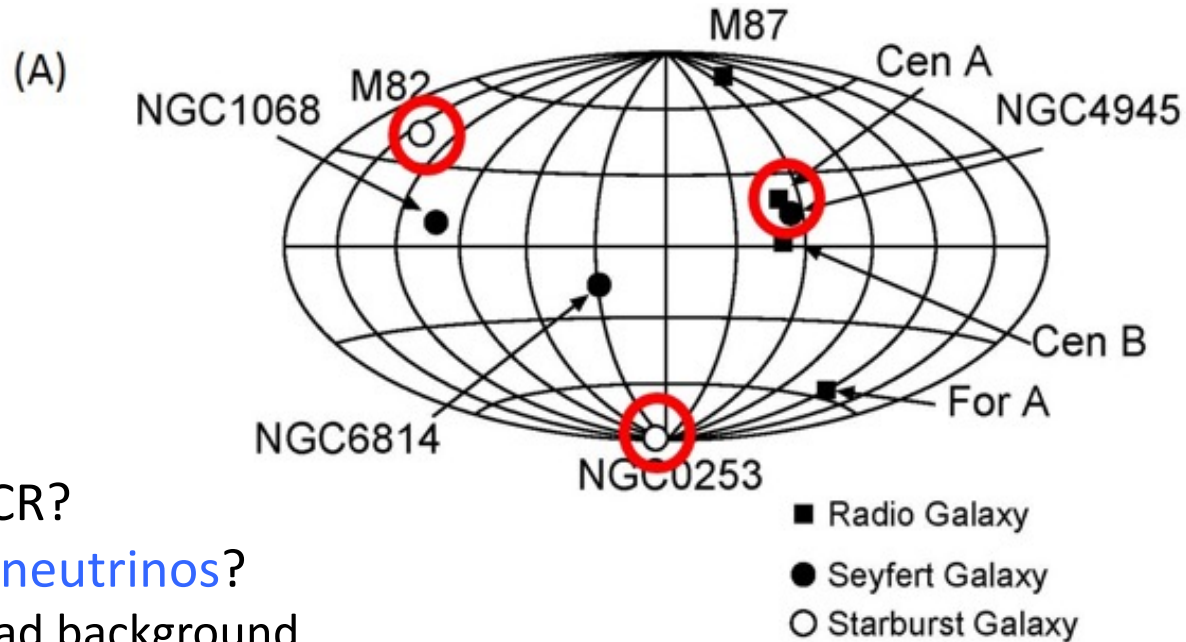


Miniquasars:
can be in our Galaxy

Ebisuzaki, Tajima
EPJ **223**, 1113(2014);
(2020)

BH Astronomy with Ultra High Energy CRs

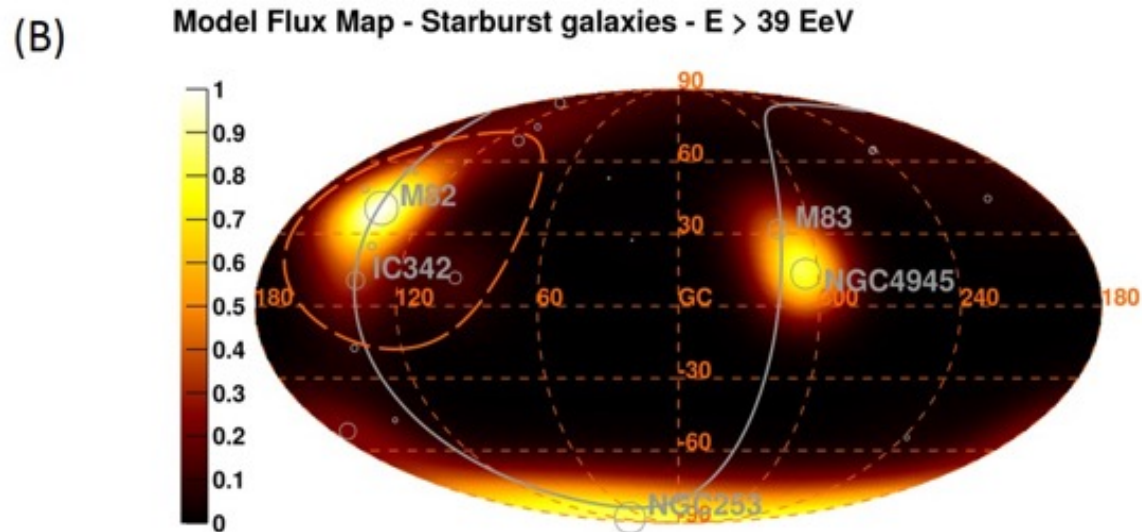
Brightest cosmic rays by wakefields



Localized UHECR?

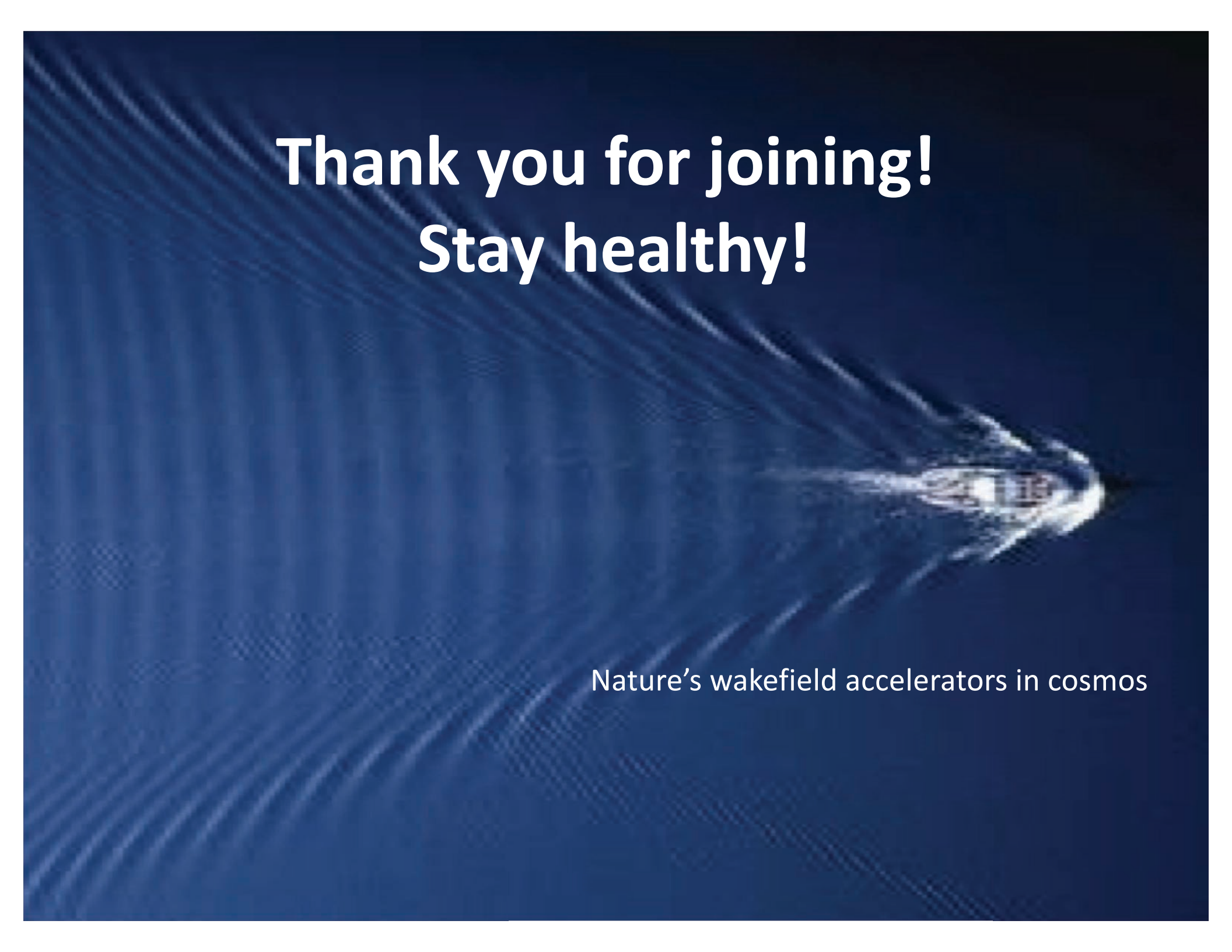
thus Localized neutrinos?

not as a spread background



Conclusions

- **Wakefield:** demonstrated ultrafast pulses, coherent collective (robust) (GeV/cm) excitable in labs (since 1994).
- Nature: more evidence of **wakefields** emerging
- **NS-NS collision:** **GW** followed by **γ -emissions**
- **Blazars:** **episodic γ -emissions** \leftarrow **wakefield** accelerated electrons \leftarrow accretion disk MRI triggered
- Nature's violent phenomena = **brightest** spots for **large** and **coherent** actions by **wakefield**
- \rightarrow **pinpointed UHECRs** (and **high energy neutrino**) arrivals
- **Gravity + plasma + B** (under certain conditions) \rightarrow plasma's theater to show **huge**, robust, highest energy, and coherent phenomenon of s.a. **wakefields**



Thank you for joining!
Stay healthy!

Nature's wakefield accelerators in cosmos